

火砕流の発生条件に対する 火口形状の影響

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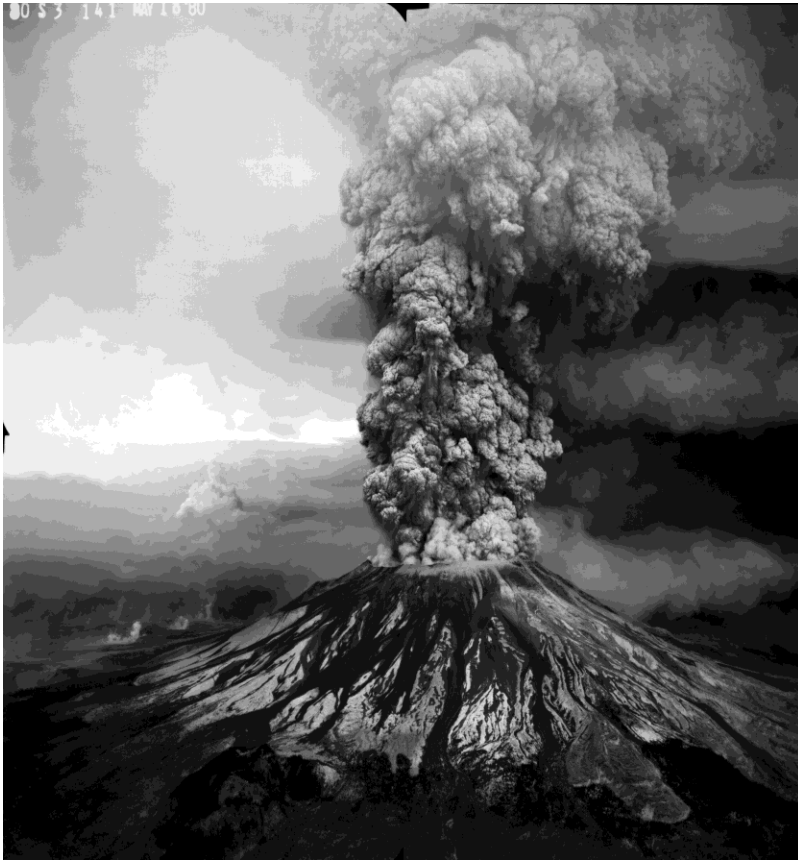
1:東大地震研

2:JAMSTEC

3:防災科研

Column collapse condition

Positively or **negatively** buoyant



Eruption column

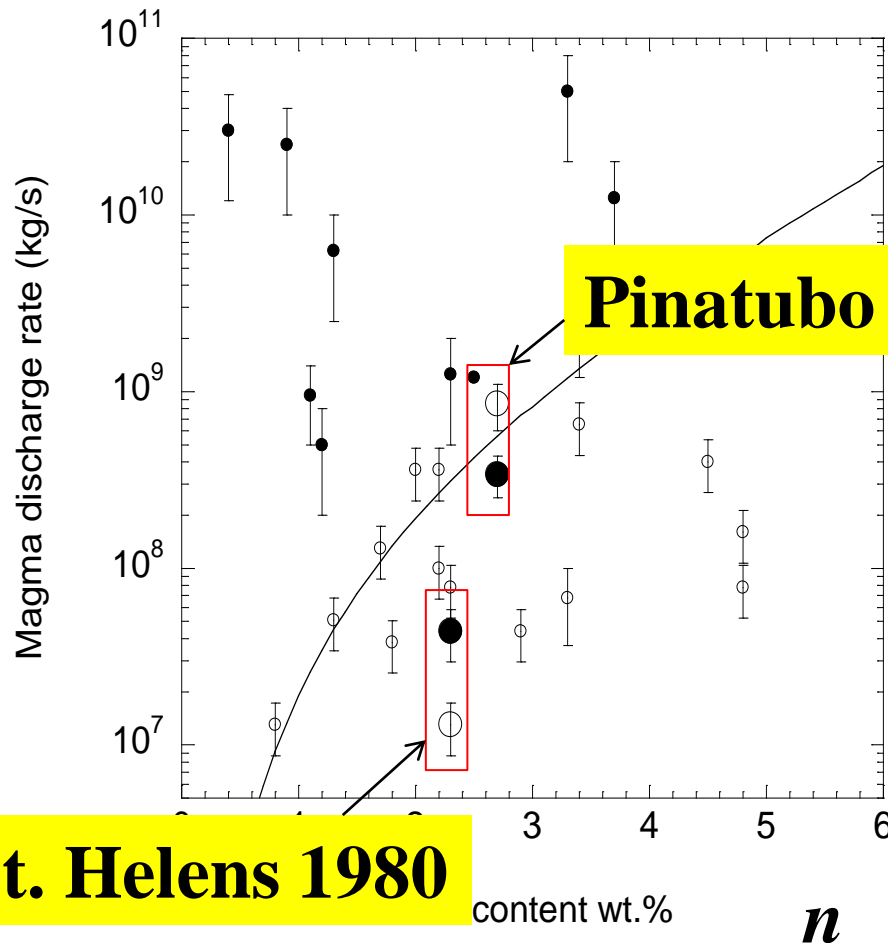
Example from the 1980 eruption
at Mt. St. Helens (USGS)



Pyroclastic flow

Previous model prediction based on 1-D steady eruption column model

Column collapse condition predicted by Carazzo et al. (2008)



Pinatubo 1991

St. Helens 1980

with the assumptions

(1) A model for entrainment by Kaminski et al. (2005)

$$(2) \quad v_a \approx 138\sqrt{n}$$

v_a : velocity at the atmospheric P just above the vent

Observations

- Pyroclastic flow
- Buoyant plume

Problems to solve

Column dynamics

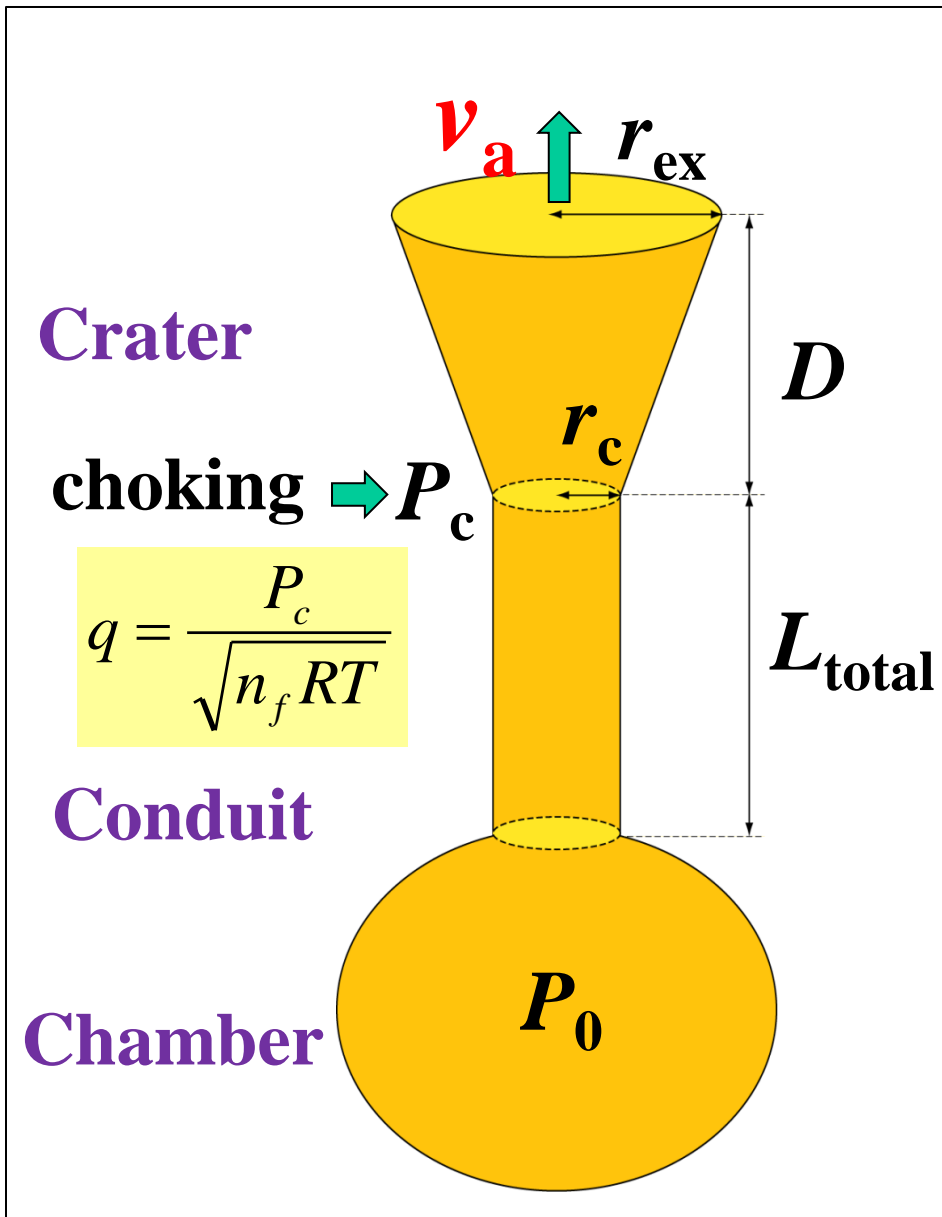
Bursik and Woods (1991)
Kaminski and Jaupart (2001)
Carazzo et al. (2008)

Flow inside crater

Woods and Bower (1995)
Ogden et al. (2008) etc.

Flow in conduit

Wilson et al. (1980)
CONFLOW (2000)
Koyaguchi (2005) etc.



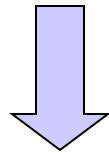
Entrainment hypothesis

[Morton et al., 1956]

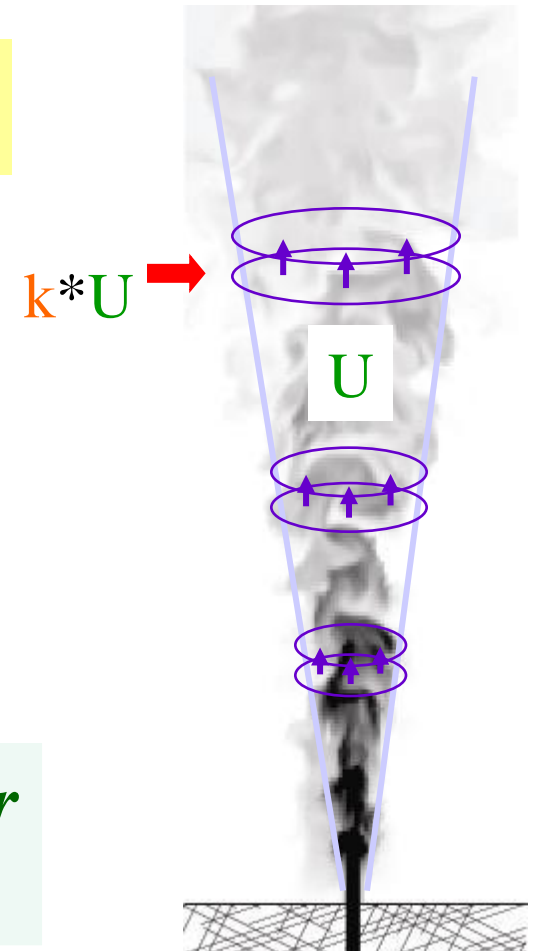
Inflow velocity = $k * \text{mean velocity}$

$k=0.09$ (e.g., Woods, 1988)

based on self-similarity of
ideal steady jets or plumes

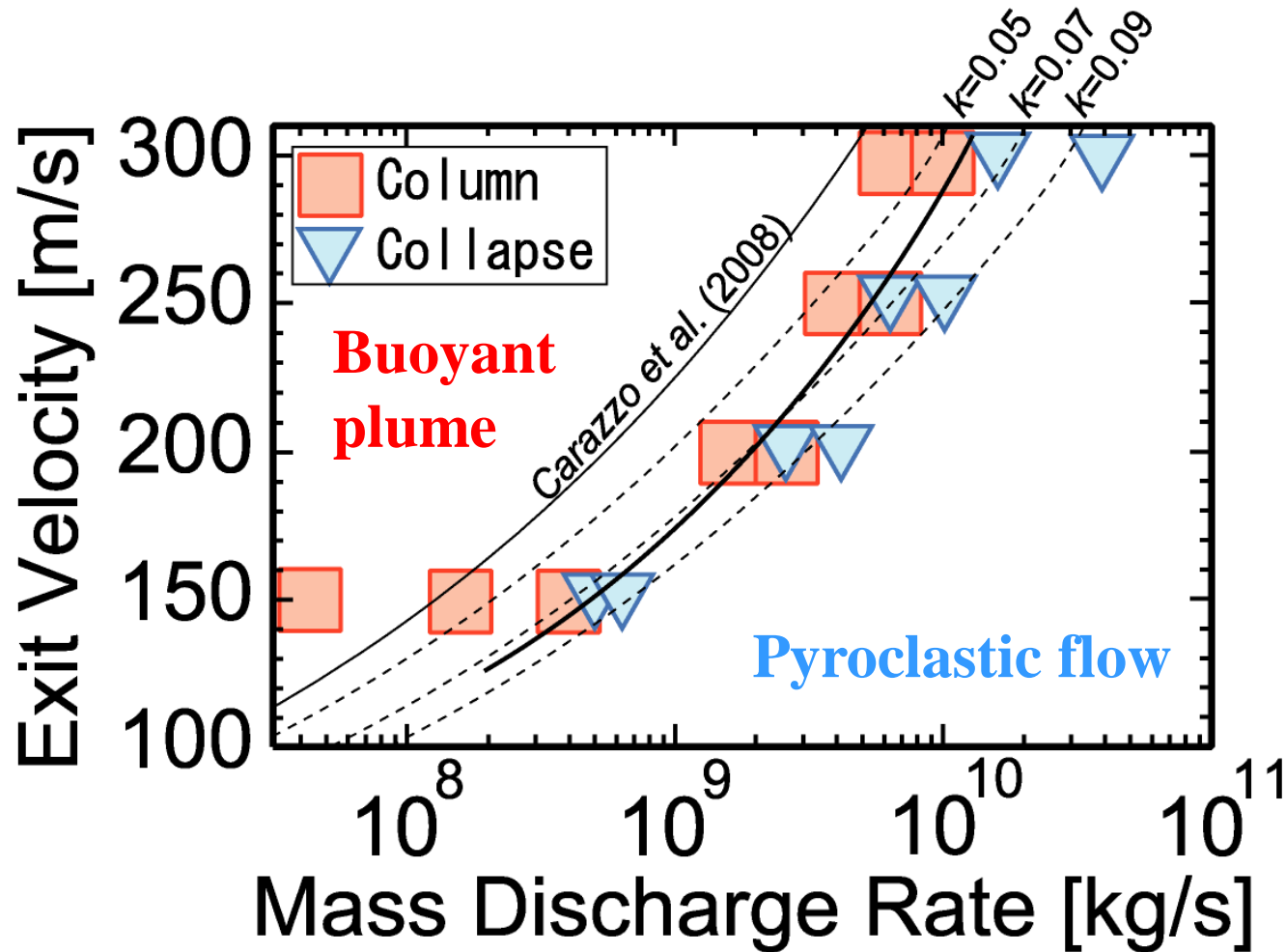


*Is this applicable to non-self-similar
eruption clouds?*



Comparison between the 1-D and 3D models

[Suzuki et al., 2005]



Problems to solve

Column dynamics

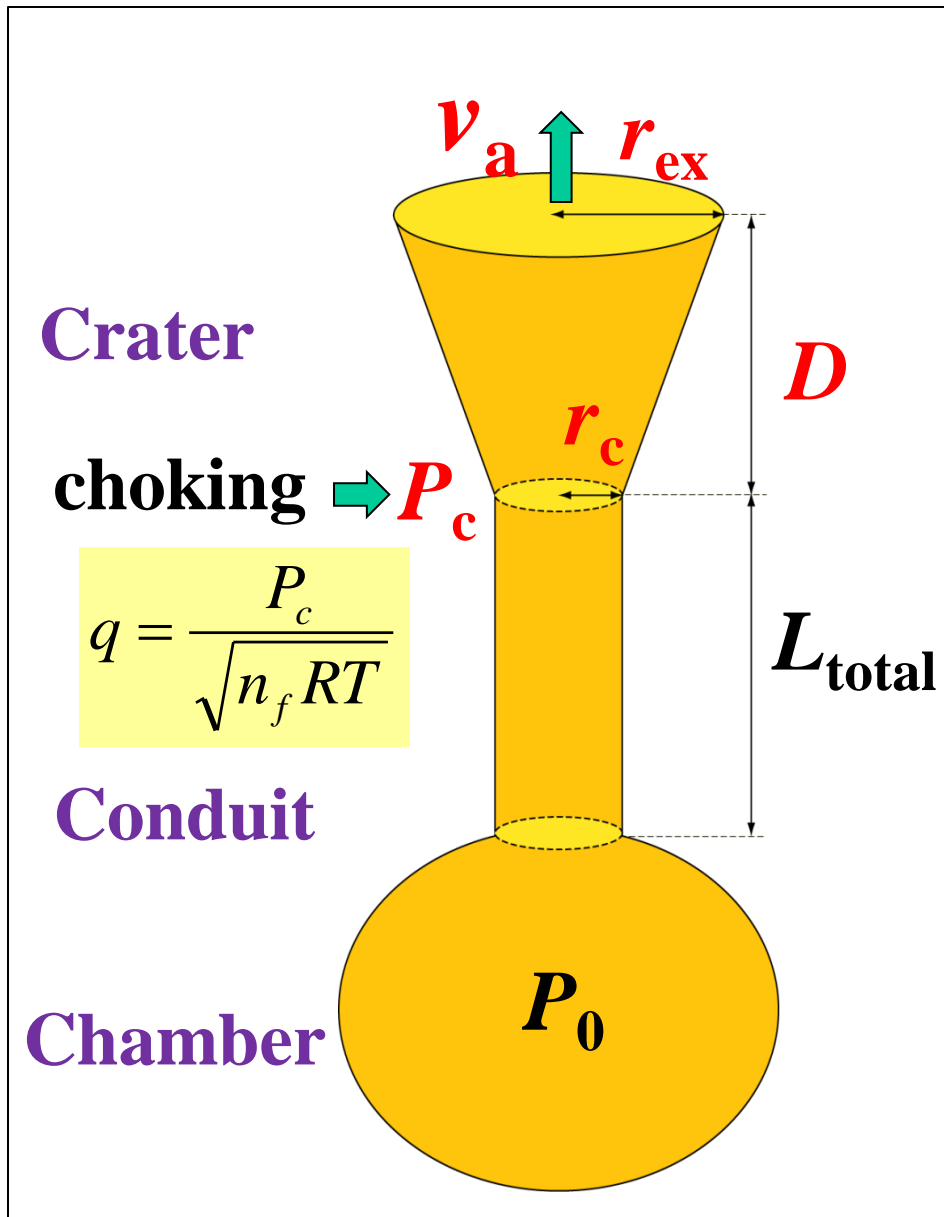
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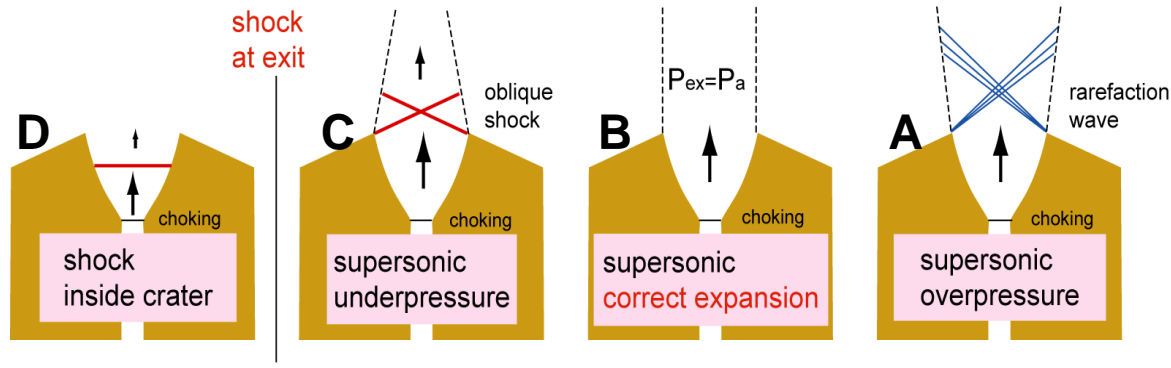
Flow in conduit

Wilson et al. (1980)
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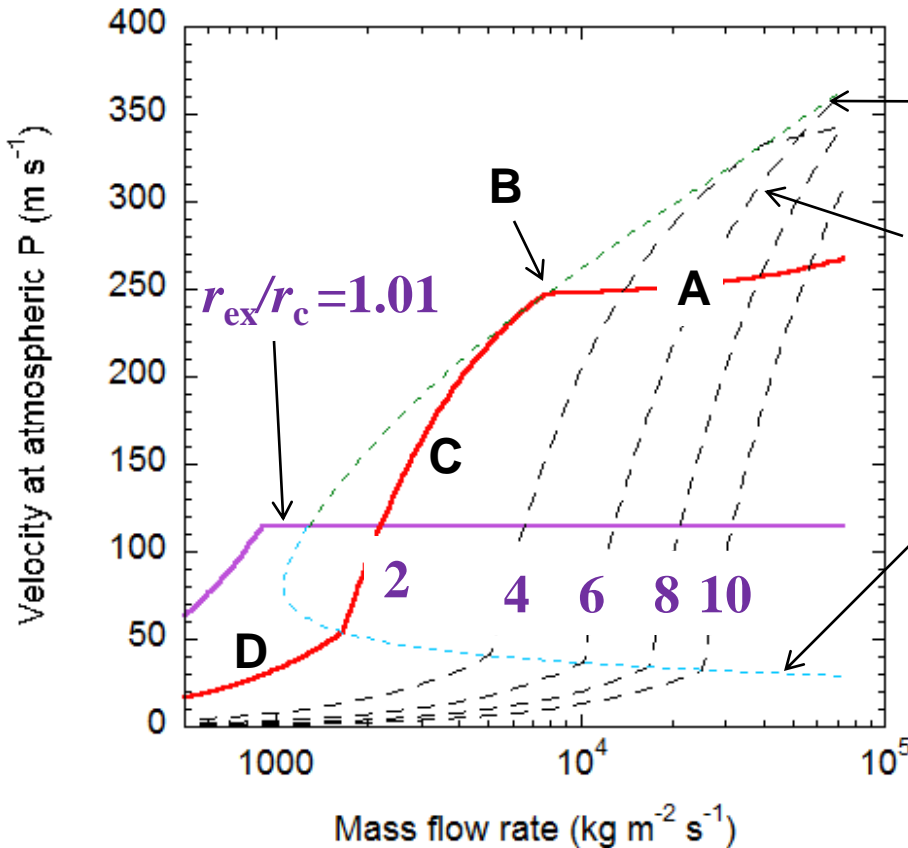
Flow inside crater

large r_{ex}/r_c
small MFR



small r_{ex}/r_c
large MFR

“ v_a ”



Correct expansion

Velocity change for given $r_{ex}/r_c (=2)$

Shock at exit

Mass flow rate (MFR)

$$q \equiv \rho_c v_c = \frac{P_c}{\sqrt{n_f RT}}$$

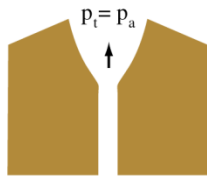
Boundaries of different flow regimes

$$\text{Eq. (2)} \quad \frac{q}{q^*} = \left(\frac{r_{ex}}{r_c} \right)^2$$

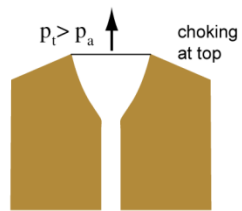
Subsonic eruption

Sonic or supersonic eruption

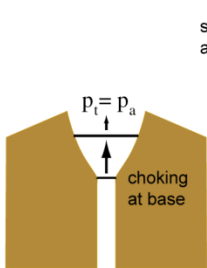
(a) subsonic throughout conduit and crater



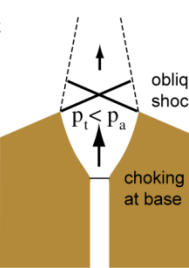
(b) free decompression



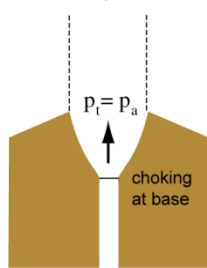
(c) shock inside crater



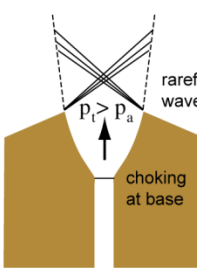
(d) overexpanded



(e) correctly expanded



(f) underexpanded



Eq. (1)

$$\frac{D}{l_{sc}} = \ln \left(\frac{r_{ex}}{r_c} \right)^2 + \frac{2V_L}{V_{Ga}} \frac{q}{q^*} \left\{ 1 - \left(\frac{r_c}{r_{ex}} \right)^2 \right\} + \frac{1}{2} \left(\frac{V_L}{V_{Ga}} \frac{q}{q^*} \right)^2 \left\{ 1 - \left(\frac{r_c}{r_{ex}} \right)^4 \right\}$$

Eq. (3)

$$\frac{1}{2} \left(\frac{q}{q^*} \right)^2 \left[\left(\frac{r_{ex}}{r_c} \right)^{-4} \left\{ \frac{V_L}{V_{Ga}} + \left(\frac{q}{q^*} \right)^2 \left(\frac{r_{ex}}{r_c} \right)^4 \right\}^2 - \left(\frac{V_L}{V_{Ga}} + \frac{q}{q^*} \right)^2 \right] + \frac{V_L}{V_{Ga}} \frac{q}{q^*} \left\{ \left(\frac{r_{ex}}{r_c} \right)^{-4} \frac{q}{q^*} - 1 \right\} + \ln \left\{ \left(\frac{r_{ex}}{r_c} \right)^{-4} \frac{q}{q^*} \right\} + \frac{D}{l_{sc}} = 0$$

Eq. (4)

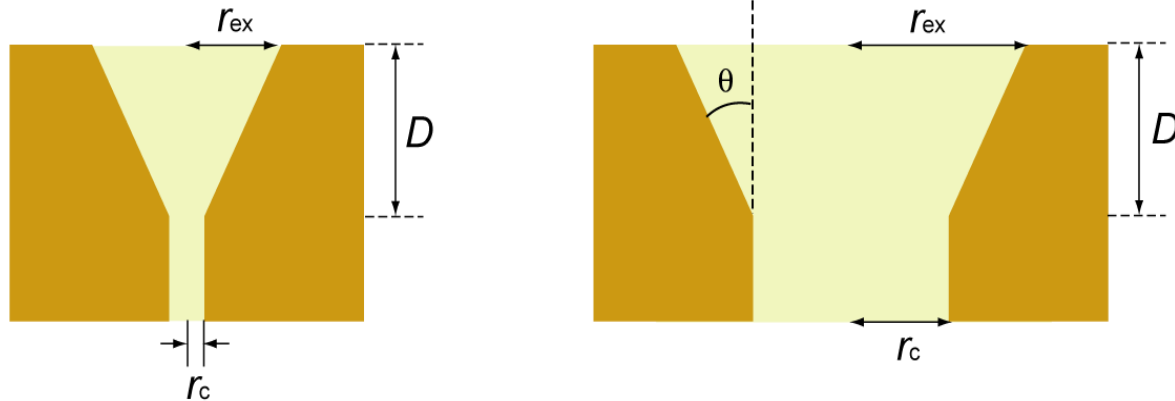
Eq. (4)

$$\frac{1}{2} \left(\frac{q}{q^*} \right)^2 \left\{ \left(\frac{r_{ex}}{r_c} \right)^{-4} \left(\frac{V_L}{V_{Ga}} + 1 \right)^2 - \left(\frac{V_L}{V_{Ga}} + \frac{q}{q^*} \right)^2 \right\} + \frac{V_L}{V_{Ga}} \left(1 - \frac{q}{q^*} \right) - \ln \left(\frac{q}{q^*} \right) + \frac{D}{l_{sc}} = 0$$

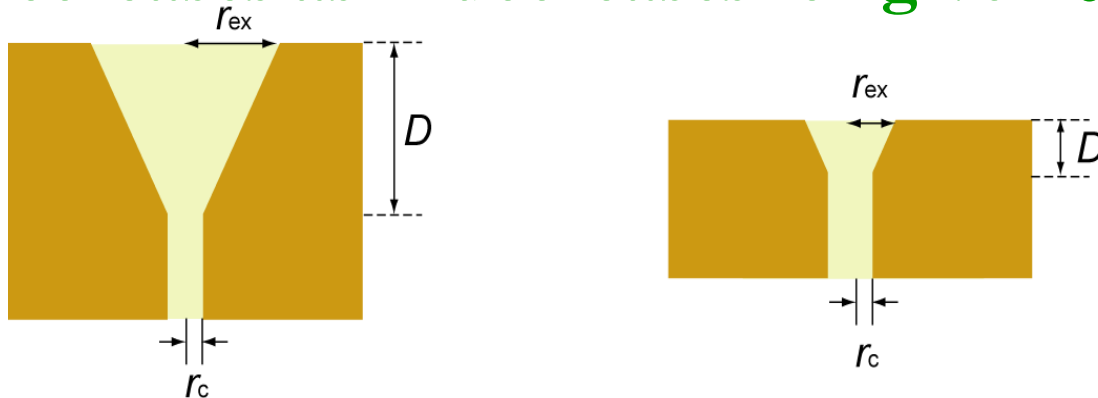
Effect of the crater shape on r_{ex}/r_c ratio

$$\frac{r_{ex}}{r_c} \approx 1 + \frac{D \tan \theta}{r_c}$$

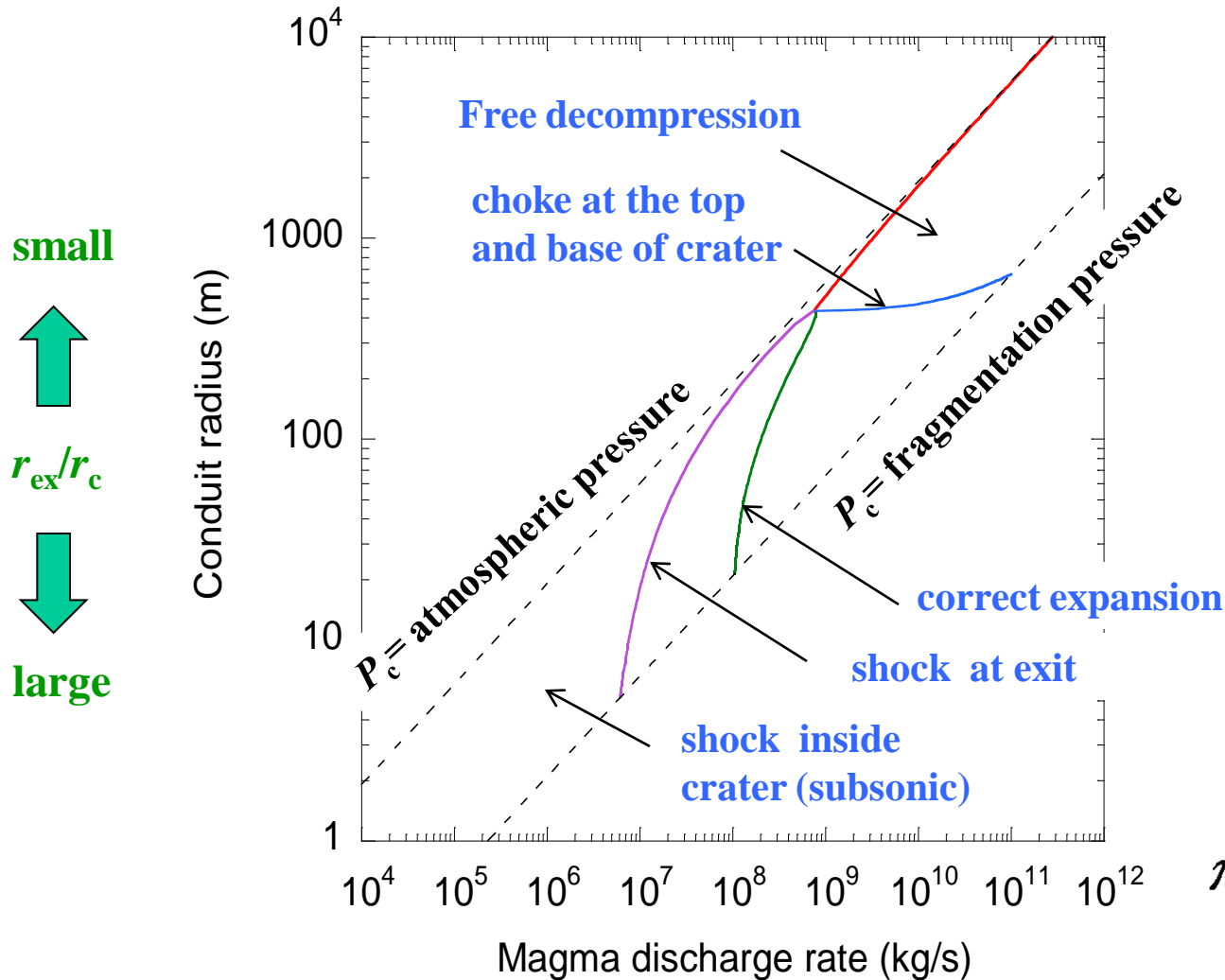
r_{ex}/r_c decreases as r_c increases for given θ .



r_{ex}/r_c decreases as D decreases for given θ .



Flow regimes on “magma discharge rate vs r_c diagram”



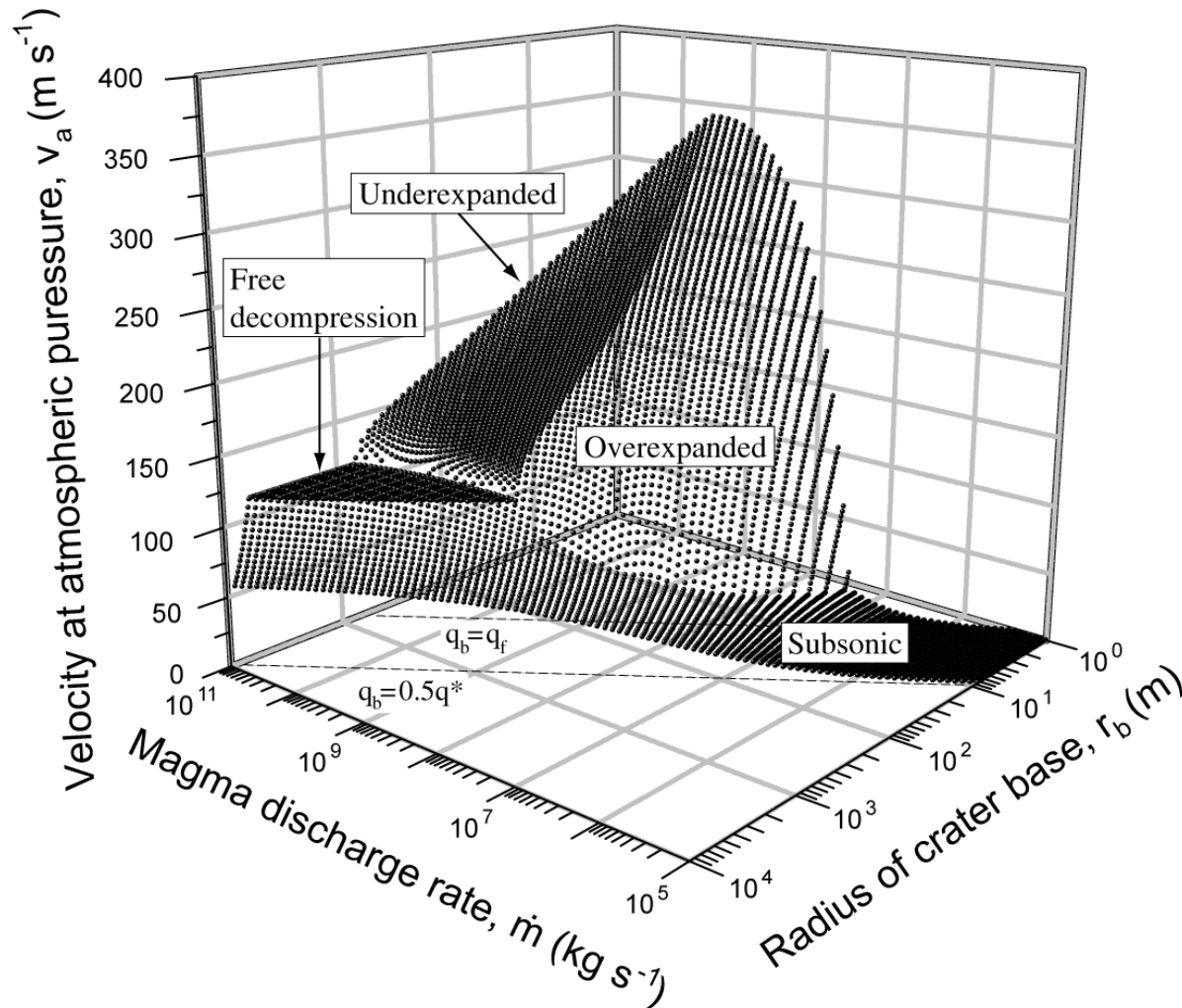
Result for
 $n_0=0.04$
 $D=500$ m
 $\theta=10^\circ$

$$\pi r_c^2 q = \frac{\pi r_c^2 P_c}{\sqrt{n_f RT}}$$

Velocity at the atmospheric P (“ v_a ”) on “magma discharge rate vs r_c diagram”

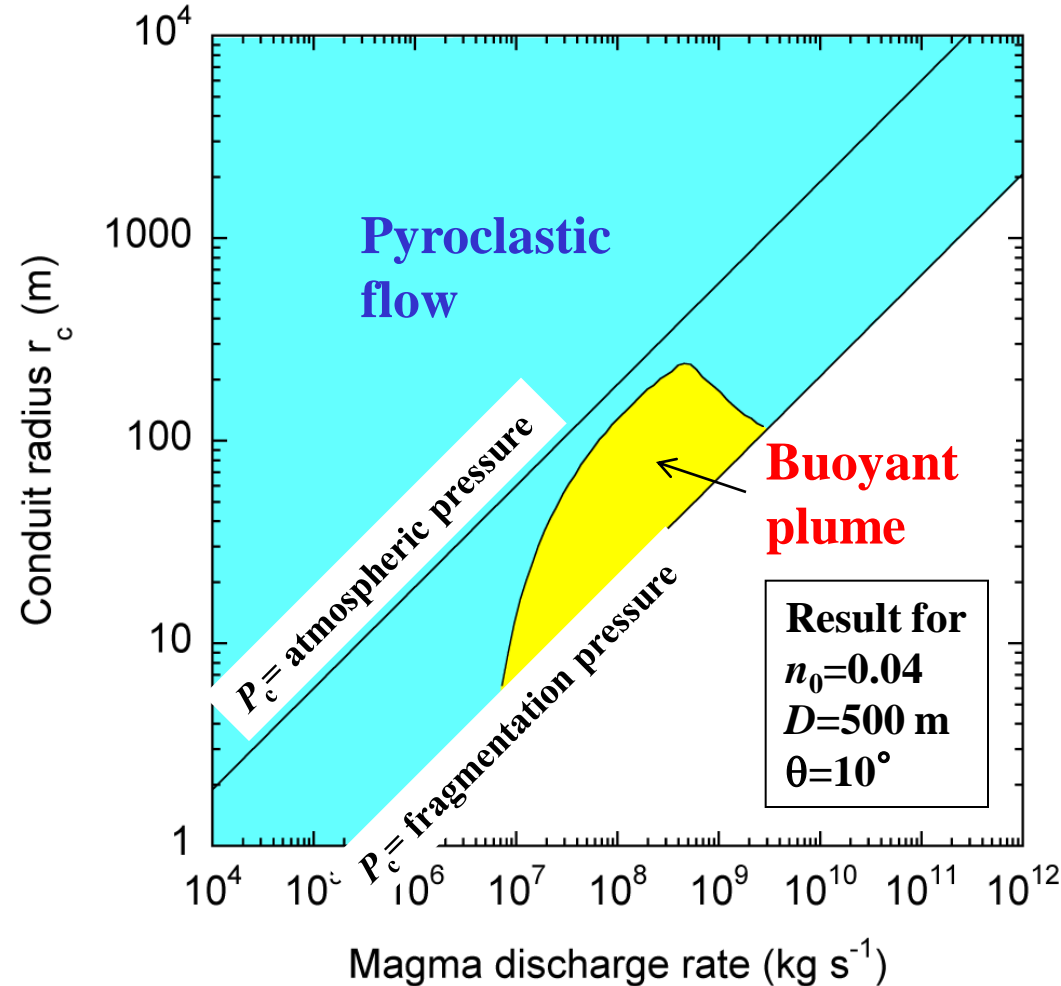
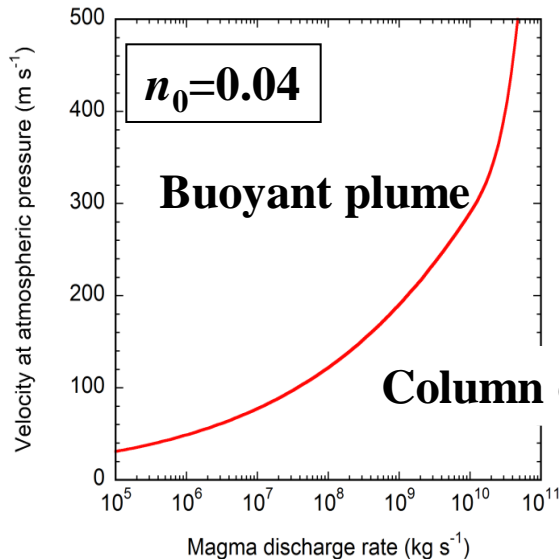
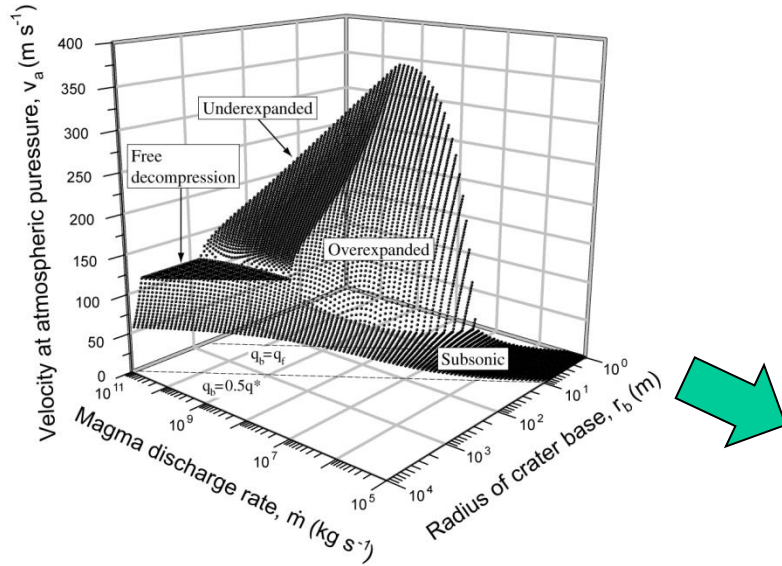
(based on Woods and Bower, 1995; Ogden et al., 2008)

“ v_a ”



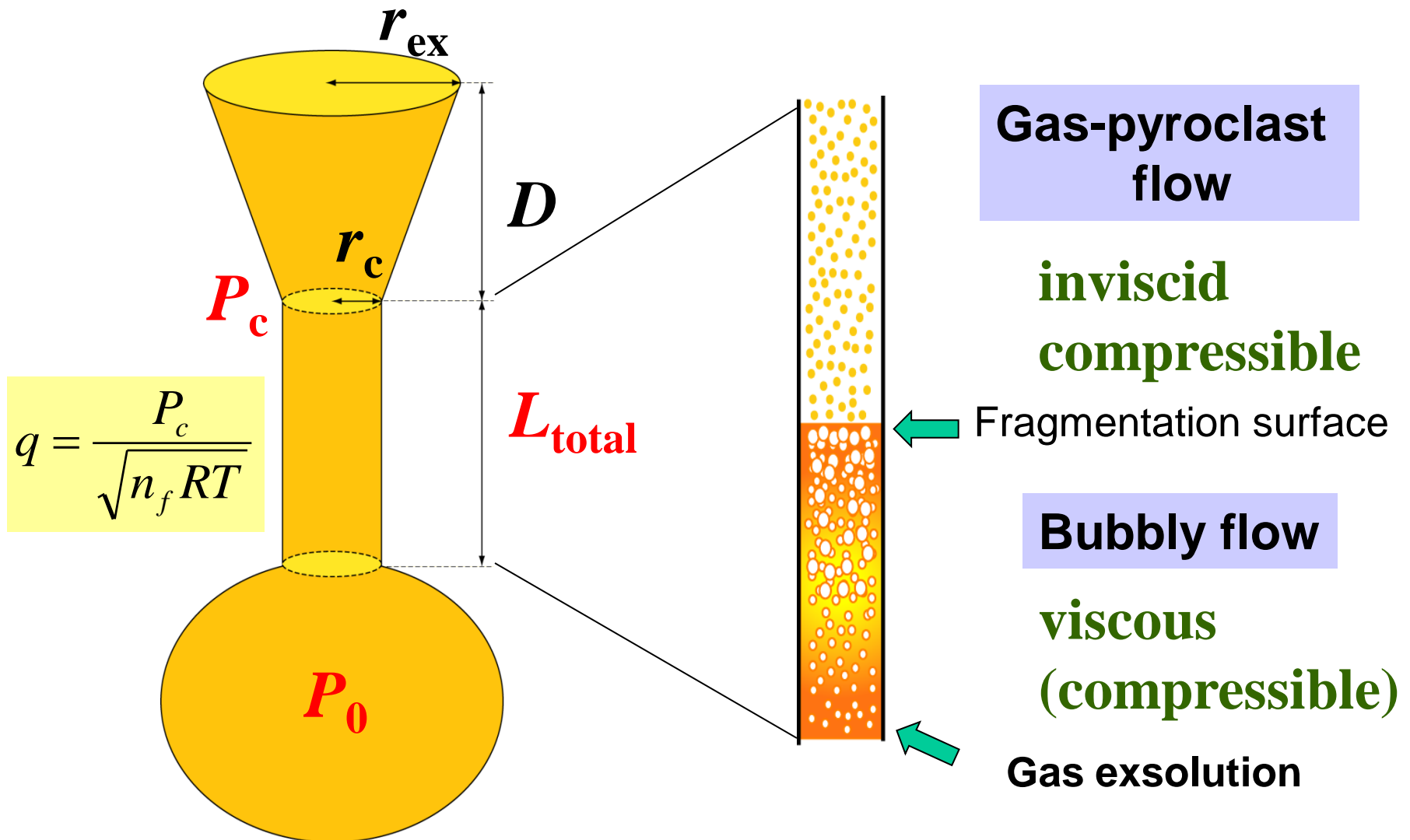
Result for
 $n_0=0.04$
 $D=500 \text{ m}$
 $\theta=10^\circ$

Proclastic-flow and buoyant plume regions on “magma discharge rate vs r_c diagram”



Mass flow rate through a conduit

1-D conduit flow model (e.g., Wilson et al., 1980; Koyaguchi, 2005)



Semi-analytical solution of 1-D steady conduit flow (Koyaguchi, 2005)

small ← $\gamma \equiv \frac{g\rho_l^2 r_c^2 \sqrt{n_f RT}}{8\eta P_f}$ → large

$\gamma \ll 1$

$\gamma \gg 1$

$$\frac{P_c}{P_f} \approx \exp\left(-\frac{gL_{total}}{n_f RT}\right)$$

Chamber depth control

$$\frac{P_c}{P_f} \approx \exp\left(-\frac{\Delta P}{\rho_l n_f RT}\right)$$

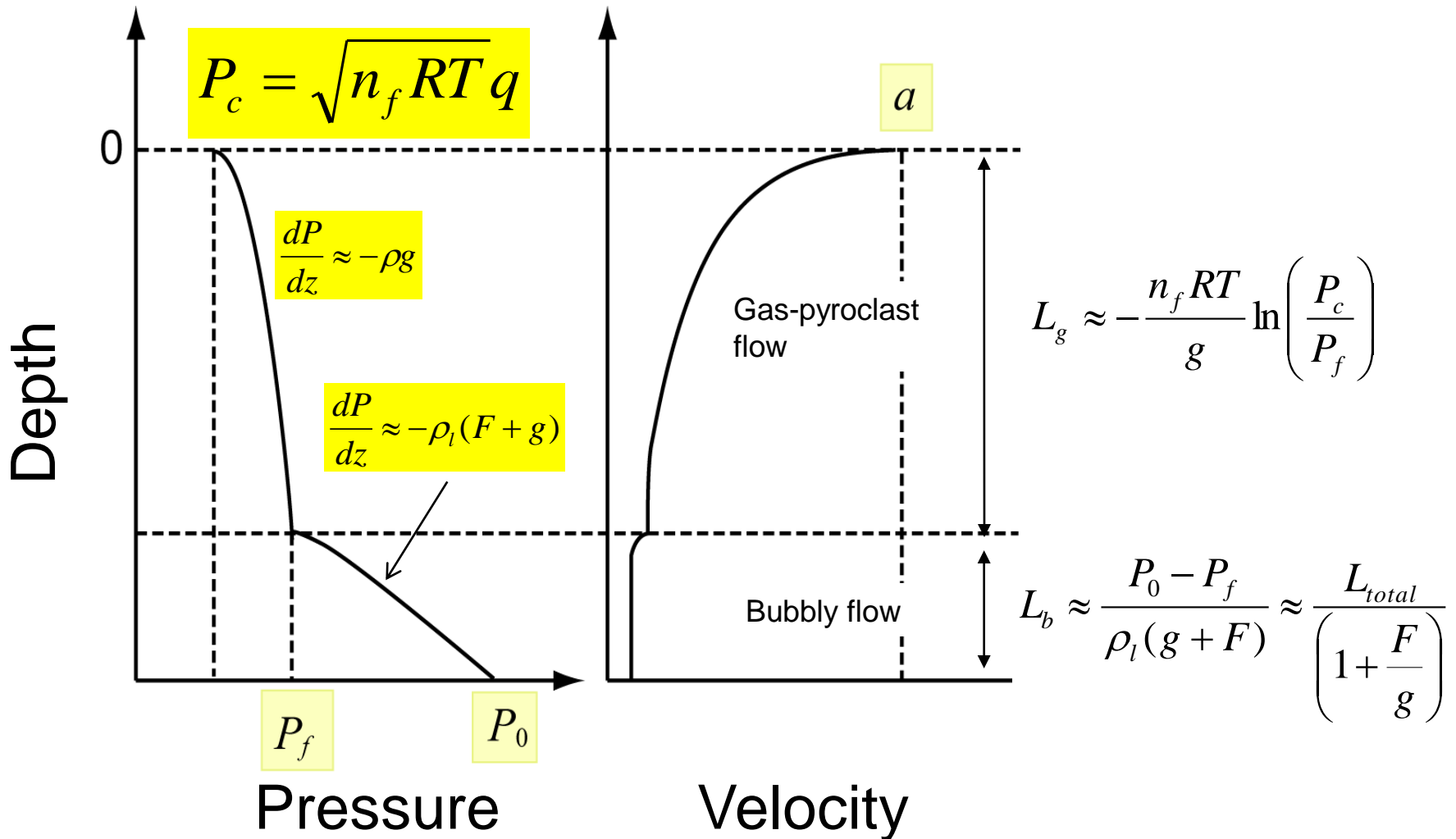
Chamber pressure control

L_{total} : chamber depth.

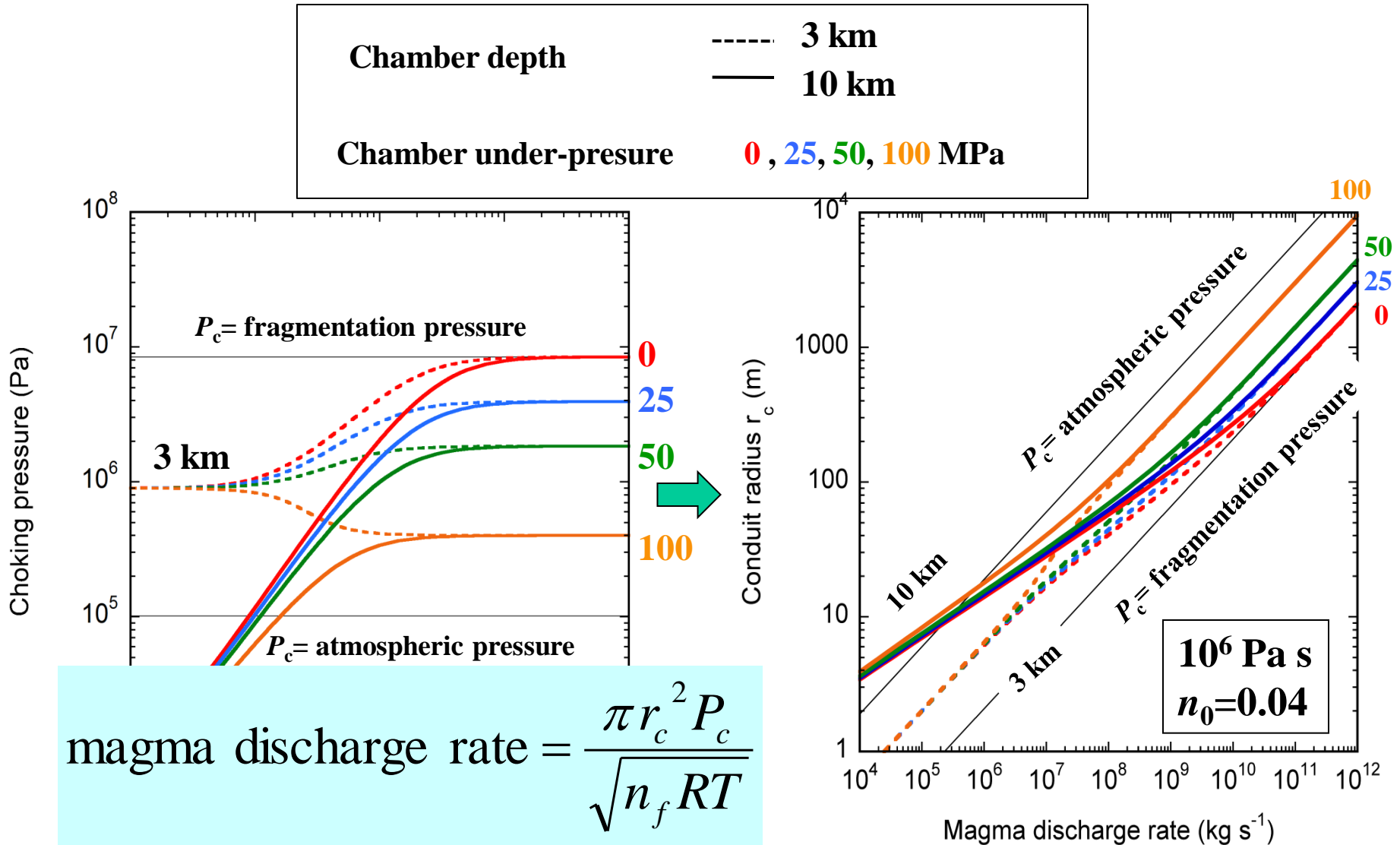
$\Delta P = P_{lith} - P_0$: degree of chamber under-pressure.

Semi-analytical solution of 1-D steady conduit flow (Koyaguchi, 2005)

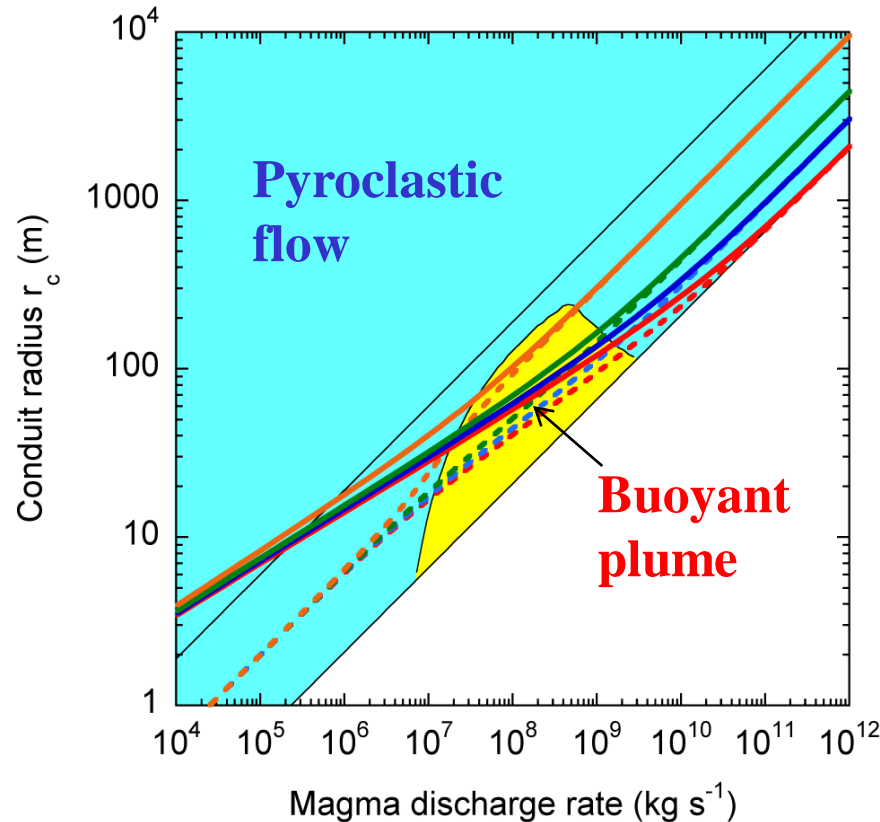
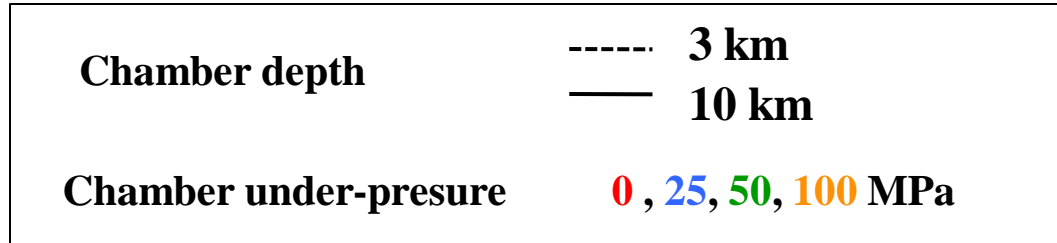
$q = \rho v$; mass flow rate



“Magma discharge rate vs r_c relationship” derived from 1-D conduit flow model



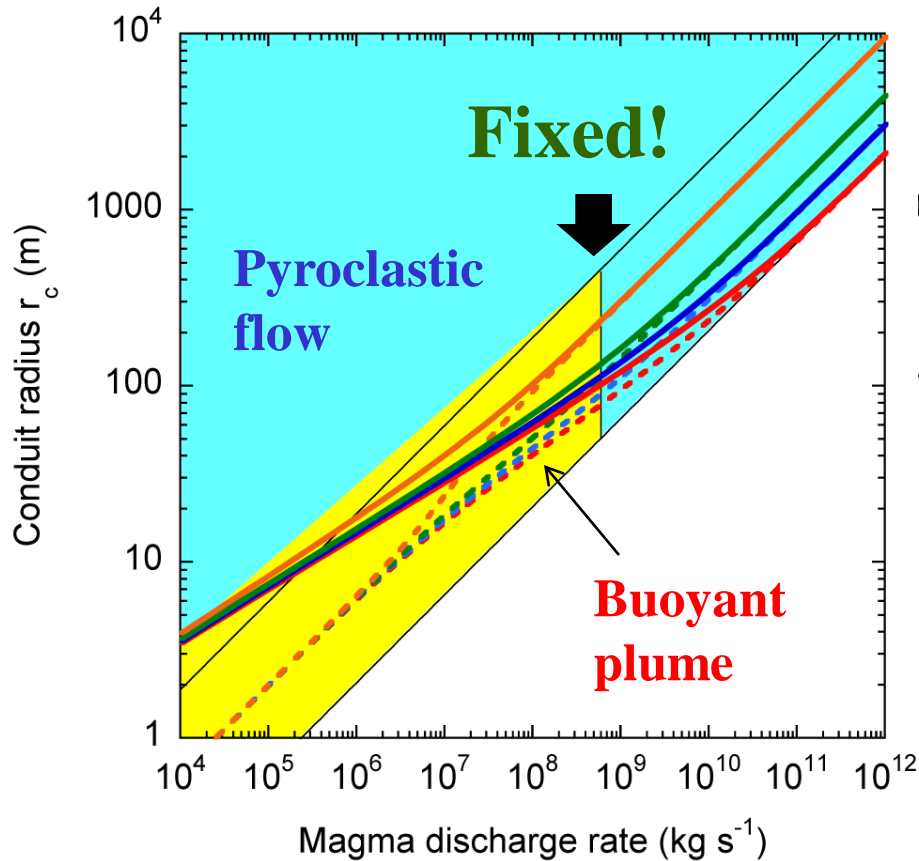
Column collapse condition on “magma discharge rate vs r_c diagram”



Result for
 $n_0=0.04$
 $D=500$ m
 $\theta=10^\circ$

Difference from the previous study

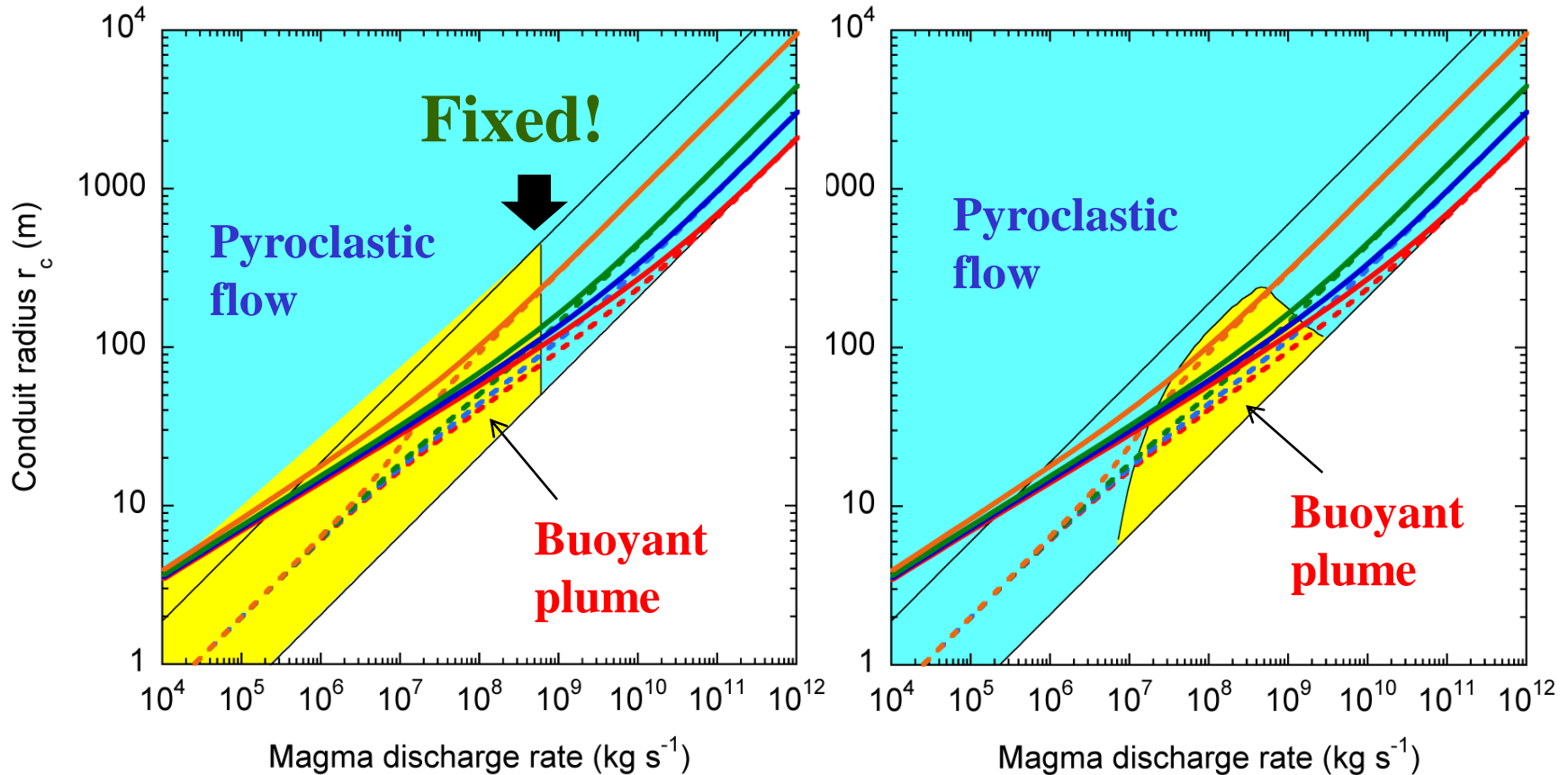
Previous study



$$v_a \approx 138\sqrt{n}$$

(e.g. Carazzo et al. 2008)

This study



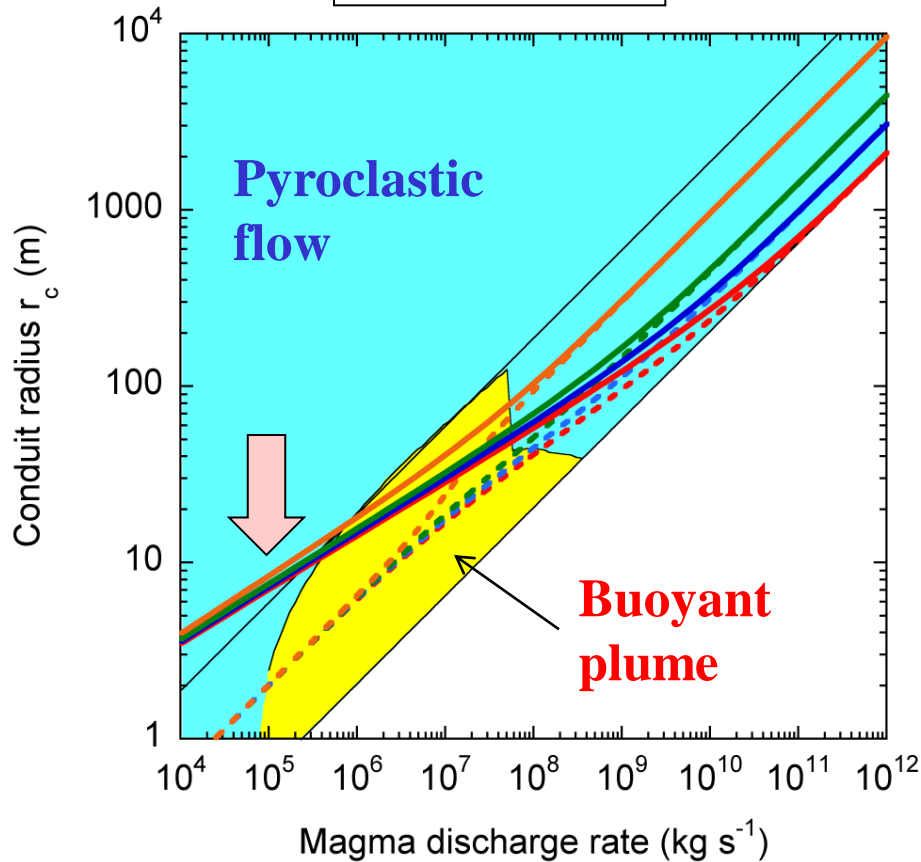
variable v_a

Dependence on crater shape

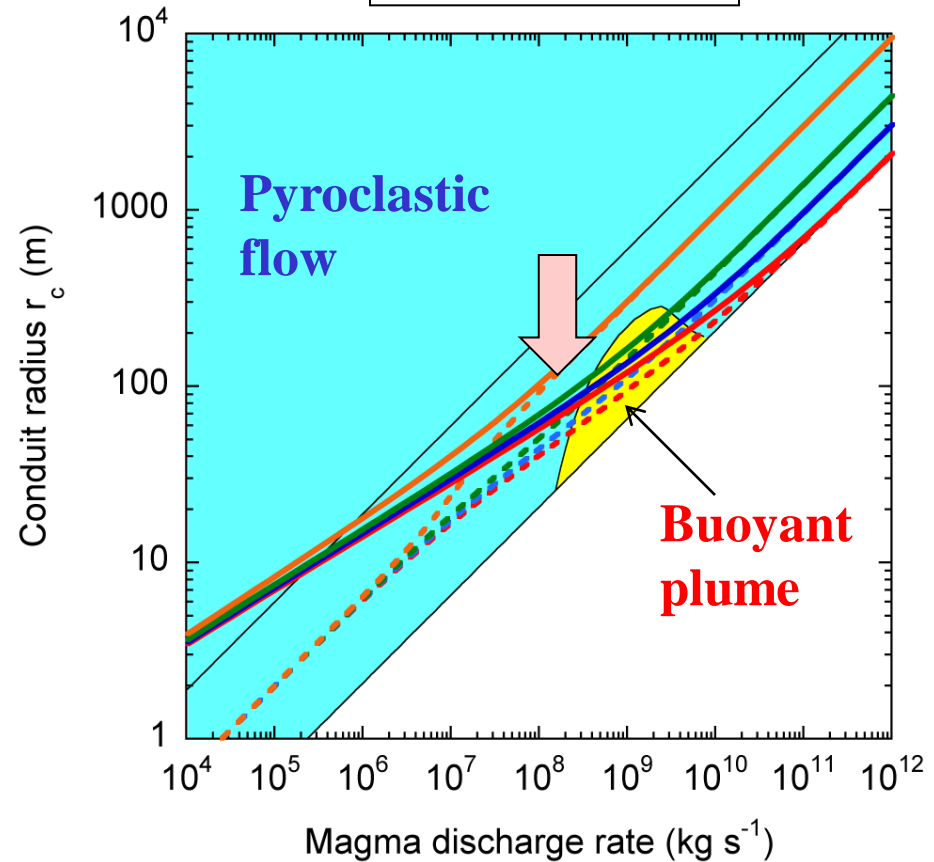
$D=500$ m
 $\theta=1^\circ$

$$\dot{m}_L \approx \frac{\pi P_a D^2 \tan^2 \theta}{\sqrt{n_f RT}}$$

$D=1000$ m
 $\theta=20^\circ$

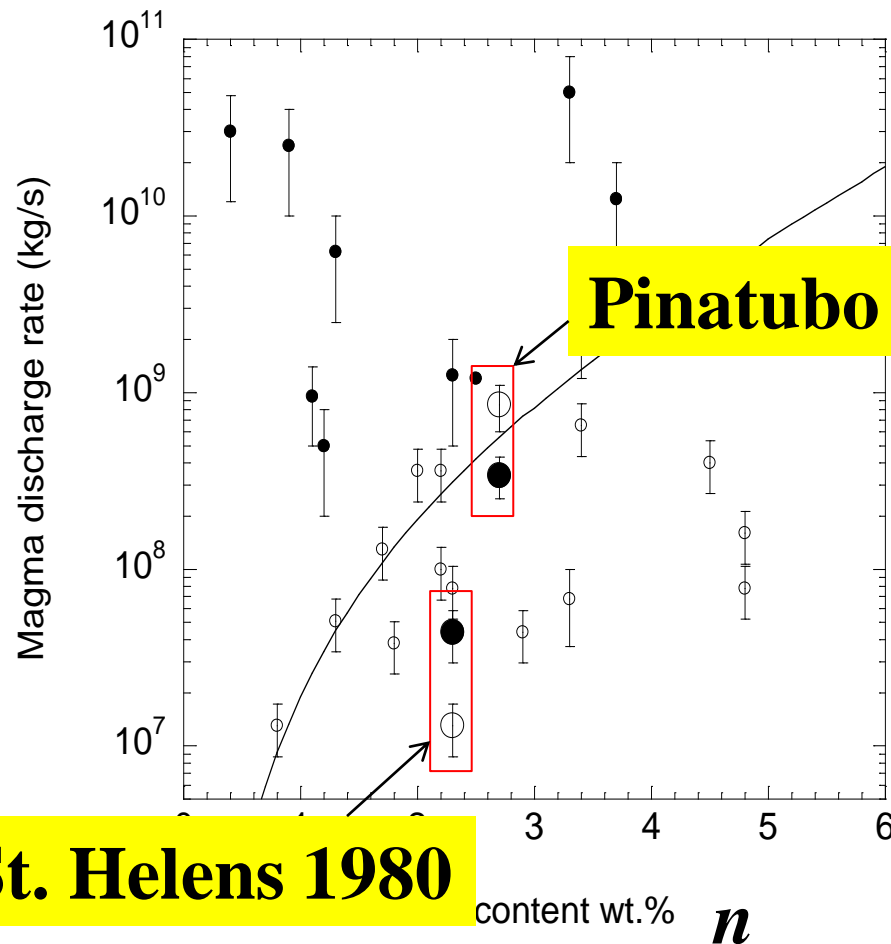


St. Helens 1980



Pinatubo 1991

Observations and previous prediction



Column collapse condition
predicted
by Carazzo et al. (2008)

Pinatubo 1991

Observations

- Pyroclastic flow
- Buoyant plume

St. Helens 1980

Conclusions and future direction

- The column collapse condition depends on entrainment coefficient, crater shape, and magma chamber pressure.
- The entrainment coefficient may be approximated by
 - $k=0.04 \sim 0.07$ near the exit
 - $k=0.10 \sim 0.15$ far from the exit.
- The effects of crater shape and magma chamber pressure on the column collapse condition can be systematically analyzed using the magma discharge rate vs r_c diagram.
- In order to confirm the present conclusions, 3-D simulations of eruption cloud are in progress, particularly focusing on more quantitative evaluation of the effects of compressible flow inside and just above the crater.