

What we found with Borexino

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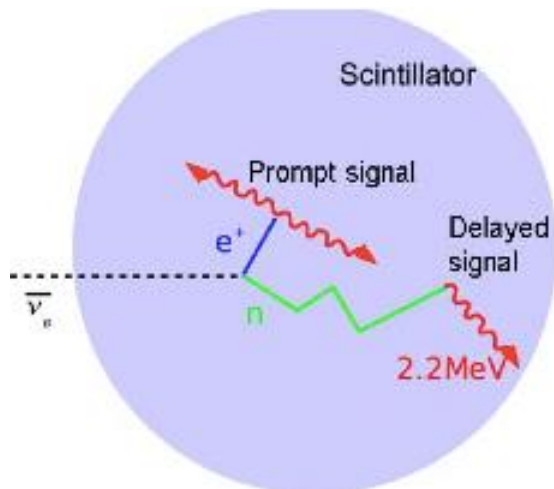
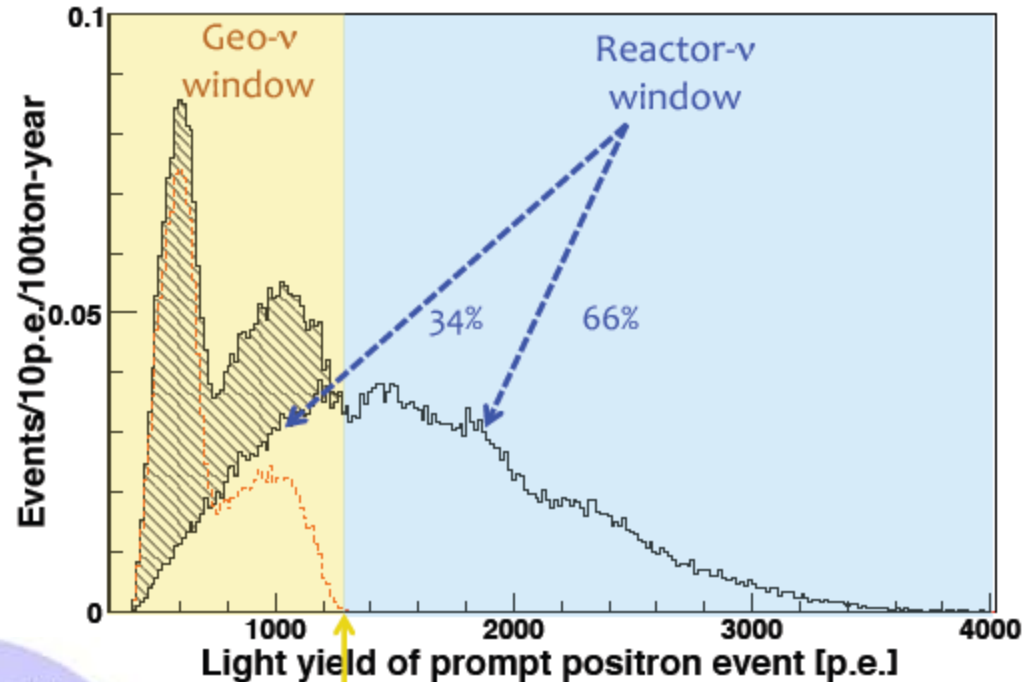
University of Houston

NMR-2013, Tokyo

Outline

- Yesterday (Jul 25): Review of Geo-neutrinos and Borexino
 - review of **Borexino** detector
 - **antineutrino** analysis
- Today (Jul 26): What we found with Borexino
 - geo-neutrino **results** (2013)
 - **implications**

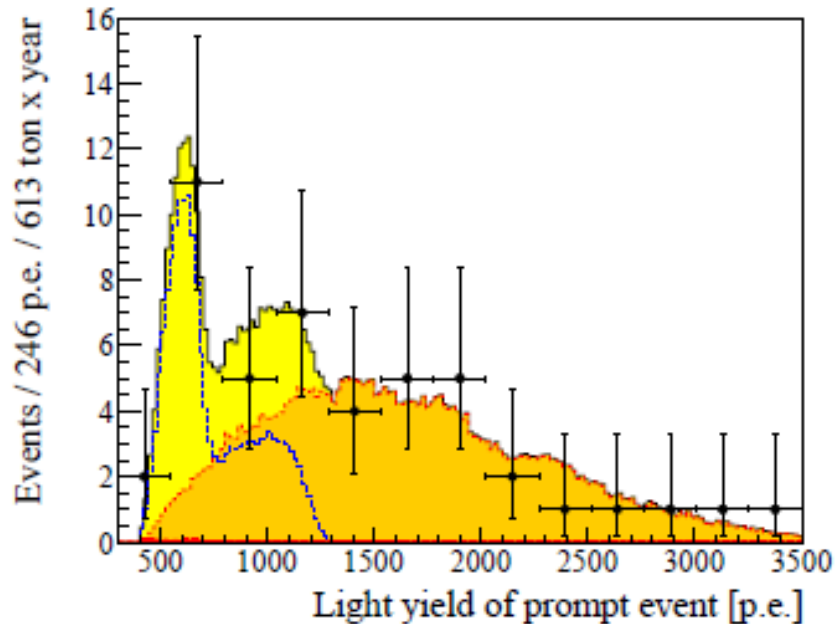
What we expect, and what the spectra means



$$E_{\text{prompt}} = E_{\bar{\nu}_e} - 0.784 \text{ MeV.}$$

In Borexino
1 Mev ~ 500 photo electrons

What we found in Borexino (1)



46 anti-nu candidates in 613 tons * year

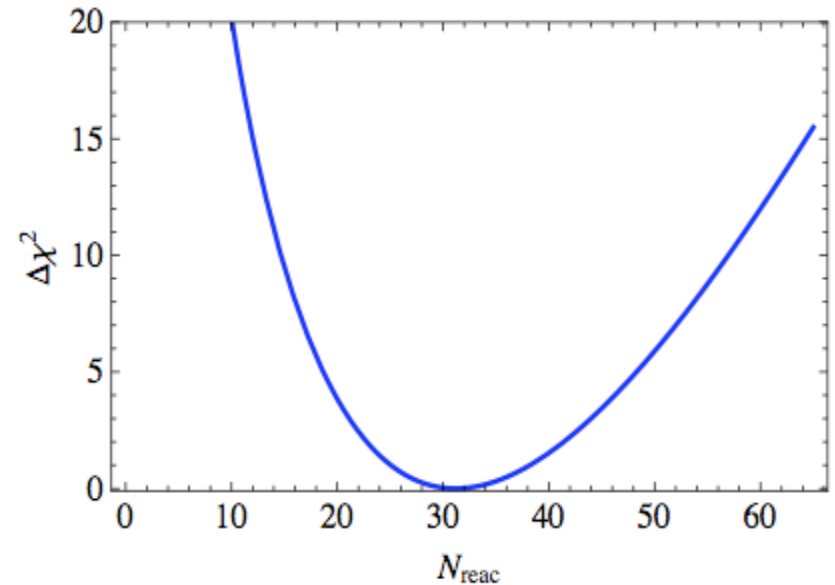
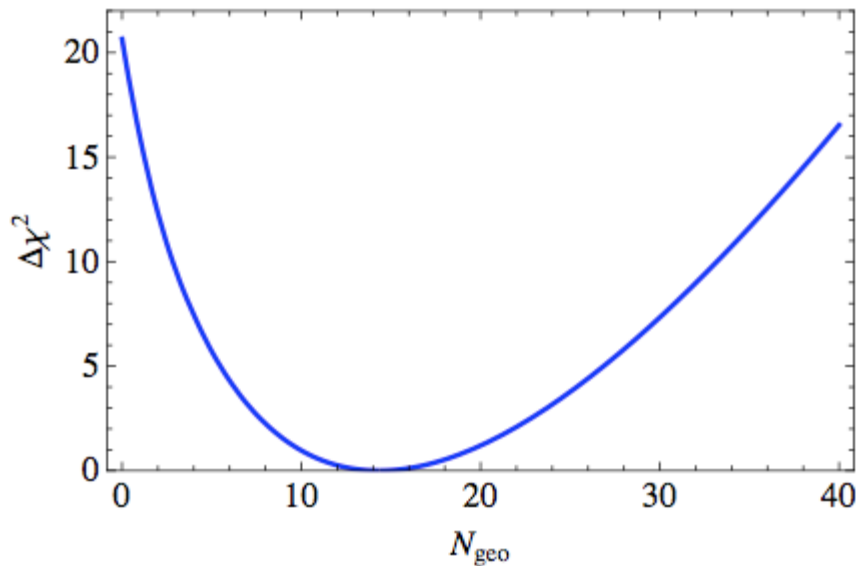
Unbinned Maximum Likelihood **fit**
to **disentangle** geo and reactors anti-nu

Fit with
Chondritic ratio
fixed: 3.9

Ngeo = (14.3 ± 4.4) events \rightarrow (38.8 ± 12) TNU

Nrea = $(31.2^{+7.0}_{-6.1})$ events \rightarrow $(84.5^{+19.3}_{-16.9})$ TNU

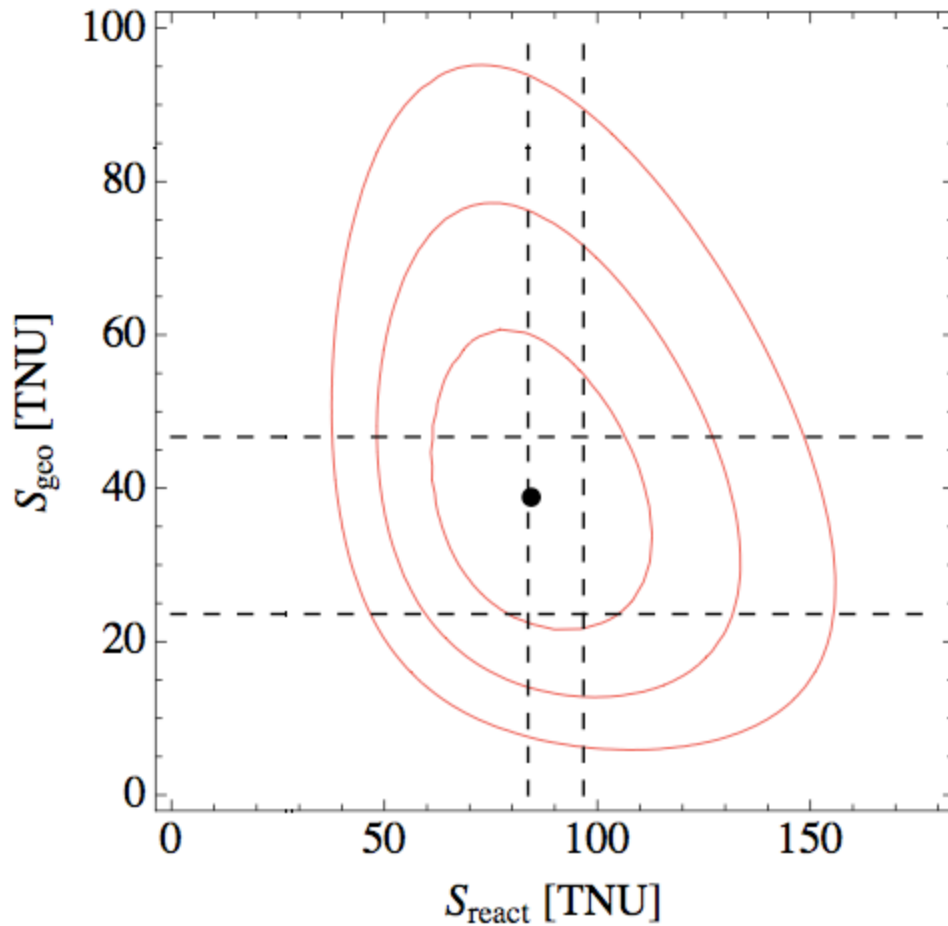
What we found in Borexino (2)



Null hypothesis for geo-neutrinos ($S_{\text{geo}} = 0$) has probability $6 \cdot 10^{-6}$

Evidence for geo-nu measurement at **4.5 σ**

What we found with Borexino (3)



Contour Plot
for geo-nu
and reactor anti-nu

1 σ expectation band
of $S_{\text{geo-nu}} = (26.3-46.6)$ TNU
for different models

1 σ expectation band of $S_{\text{reactor}} : (83.2-97.3)$ TNU

Mantle contribution

$$S_{\text{geo}} = S(\text{Crust}) + S(\text{Mantle})$$

$$S^{\text{BX}}_{\text{geo}}(\text{Total}) = (38.8 \pm 12.0) \text{ TNU}$$

$$S^{\text{BX}}_{\text{geo}}(\text{Crust}) = (23.4 \pm 2.8) \text{ TNU}$$

M. Coltorti et al., Geo.Cosm. Acta 75(2011)2271

Y. Huang et al., arXiv:1301.0365

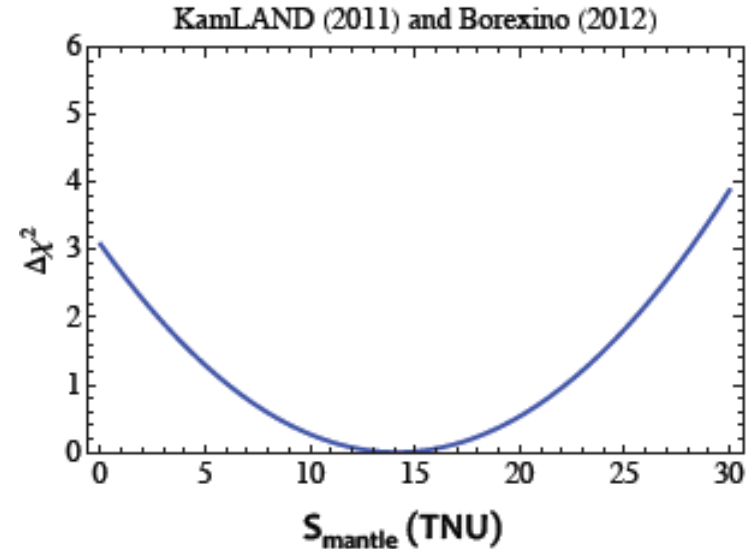
$$S^{\text{BX}}(\text{mantle}) = (15.4 \pm 12.3) \text{ TNU}$$

Combined analysis with KamLAND

$$S^{\text{KL}}_{\text{geo}}(\text{Total}) = (38.3^{+10.3}_{-9.9}) \text{ TNU}$$

$$S^{\text{KL}}_{\text{geo}}(\text{Crust}) = (25 \pm 2) \text{ TNU}$$

A. Gando et al. (KamLAND Collab.), Nature Geosci. 4 (2011) 647



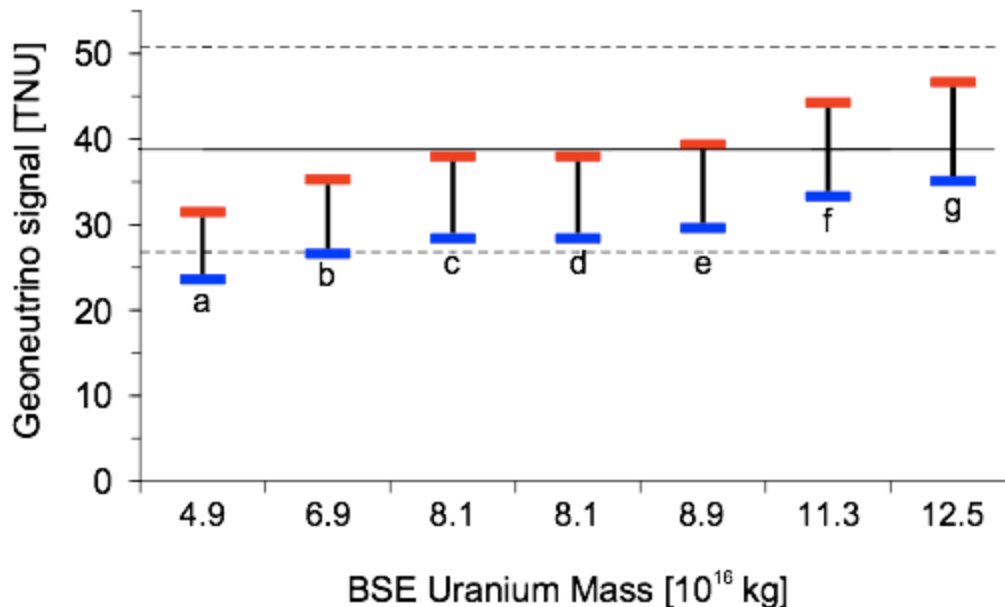
$$S(\text{Mantle}) = (14.1 \pm 8.1) \text{ TNU}$$

Comparison with BSE models

Total S_{geo} [TNU]		Model
low	high	
35.1	46.64	Turcotte & Schubert 2002 (g)
33.3	44.24	Anderson 2007 (f)
29.6	39.34	Palme & O'Neil 2003 (e)
28.4	37.94	Allegre et al. 1995 (d)
28.4	37.94	Mc Donough & Sun 1995 (c)
26.6	35.24	Lyubetskaya & Korenaga 2007 (b)
23.6	31.44	Javoy et al. 2010 (a)

High = homogeneous mantle, crustal signal + 1 σ error

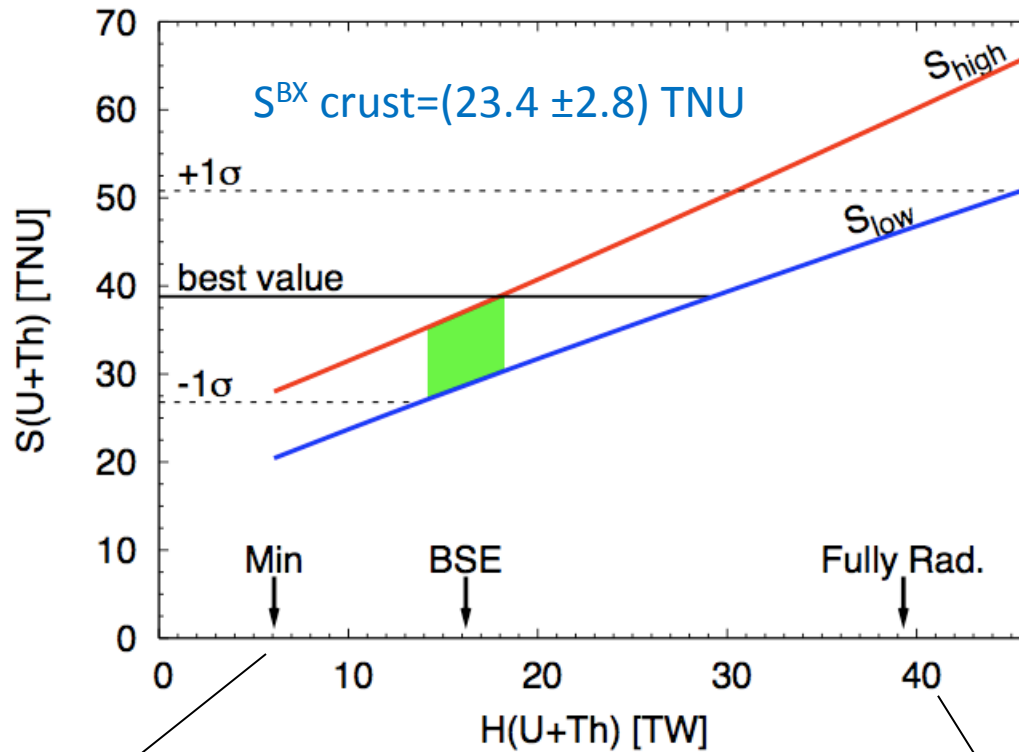
Low = HPE's concentrated close to the core-mantle boundary, crustal signal -1 σ error



1 σ band

BSE models in agreement with geo-nu measurement

Earth radiogenic power



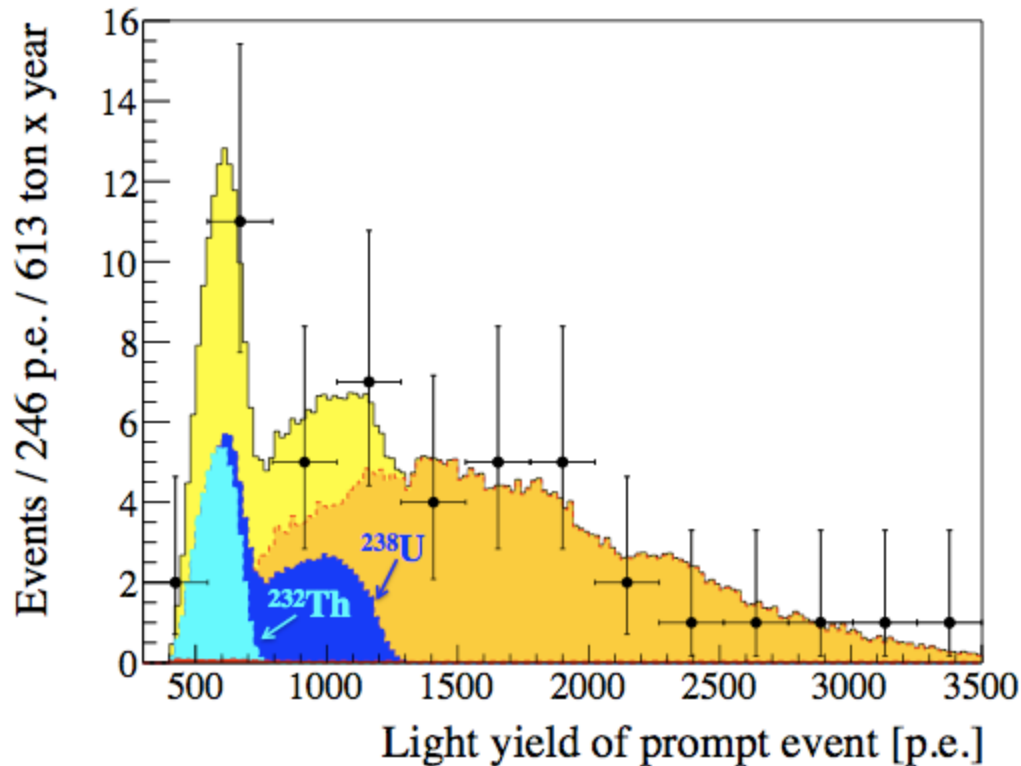
Red line: ROC+LOC +1 σ ,
mantle homogeneous

Blue line: ROC+LOC -1 σ ,
HPE's at mantle-core
boundary

Min = only crustal HPE's

Full Rad: Tot Heat = (47 ± 2) TW $\Rightarrow H_U + H_{Th} = 39.3$ TW

What we found with Borexino (4)



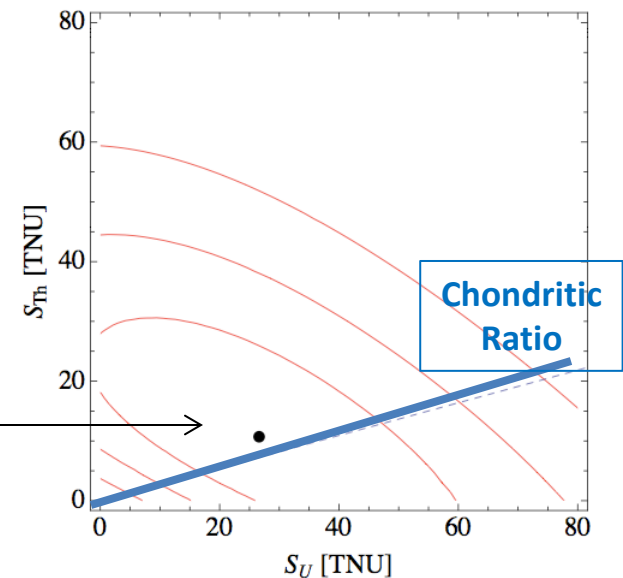
$S_U = (26.5 \pm 19.5) \text{ TNU}$
 $S_{\text{Th}} = (10.6 \pm 12.7) \text{ TNU}$

Fit with separate (free)
U and **Th** contribution

$N_U = 9.8 \pm 7.2$ events

$N_{\text{Th}} = 3.9 \pm 4.7$ events

$N_{\text{react}} = 31.7^{+7.2}_{-6.3}$ events



Summary and Outlook

New **Borexino** result on geo- ν fluxes measured at **LNGS**,
fiducial exposure of 613 ton*year ($3.69 \cdot 10^{31}$ N protons)

Measured **(14.3 \pm 4.3) geo- ν , (38.8 \pm 12) TNU**

BSE models in agreement with Geo- ν measurement

Signal for **mantle** deduced,
combined analysis with **KamLand**
S(mantle) = (14.1 \pm 8.1) TNU

Spectroscopic **measurement** of U and Th,
obtained best value for **Th/U ratio** close to the chondritic one

Just one more step to attack models!