

グローバル地震観測

地震研究所
海半球観測研究センター
川勝 均

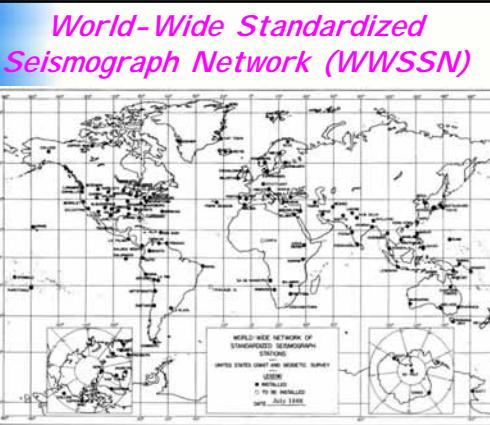
グローバル地震観測網（定常 / 機動）
震源解析（CMT）
地震波トモグラフィー
マントル
遷移層
CMB
内核
次世代観測網の展開
青い地球の地震学



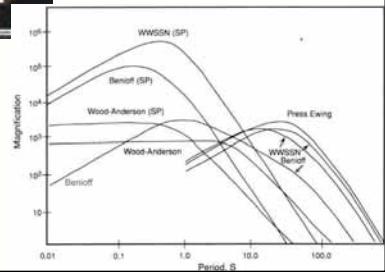
Global Seismic Network - 0th generation -



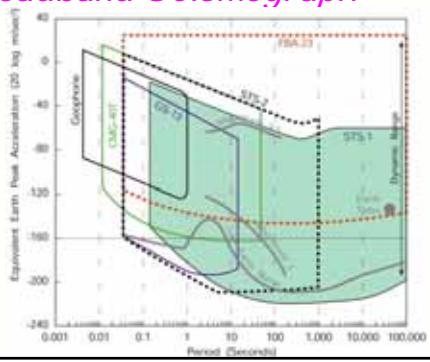
Figure 1. World wide network of stations reporting data to Milne at Shide, around 1910.



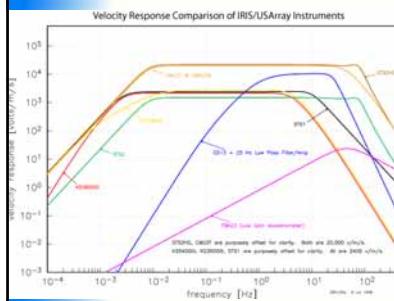
WWSSN



-Current Generation- "Broadband Seismograph"

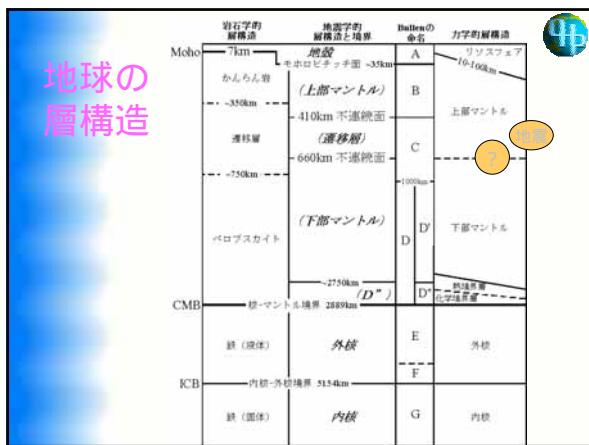
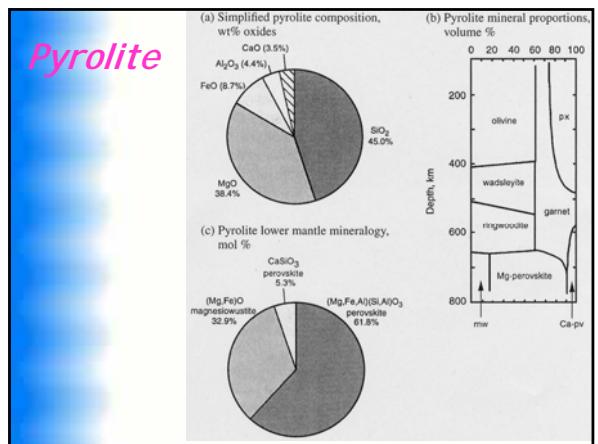
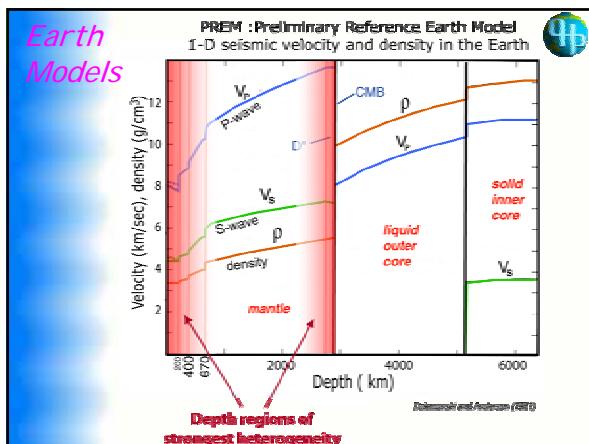
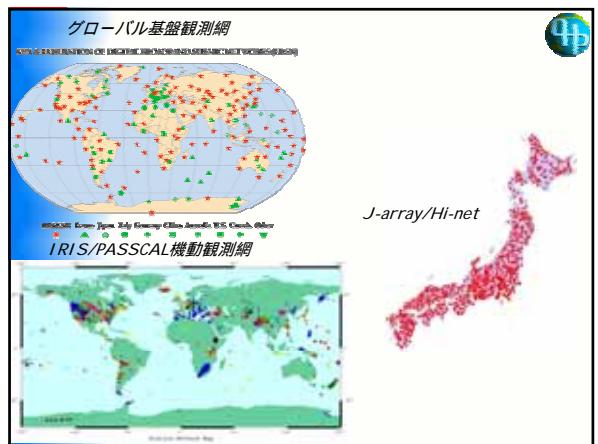
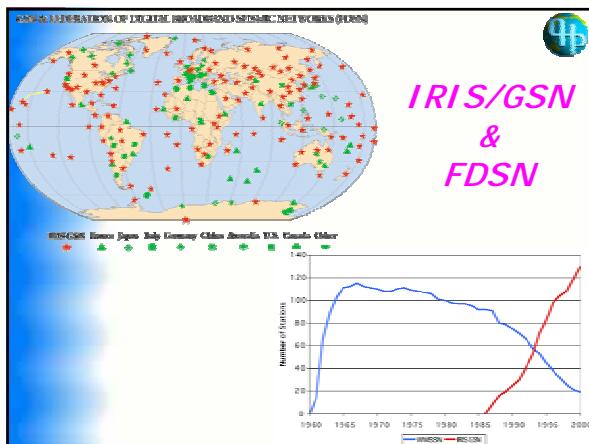


Broadband Seismograph

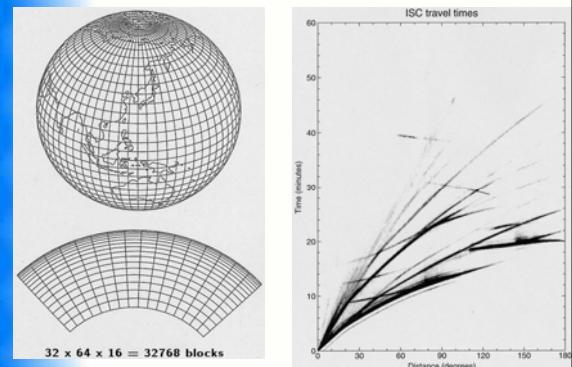


"STS1"

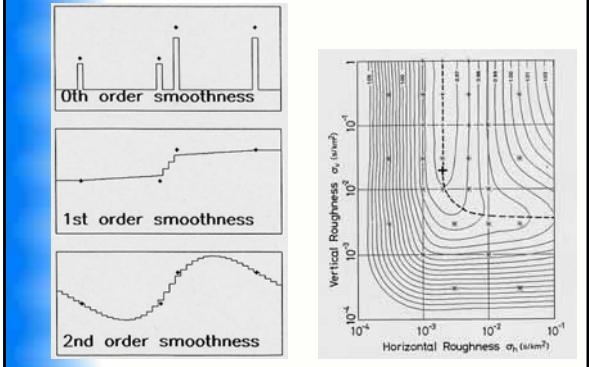




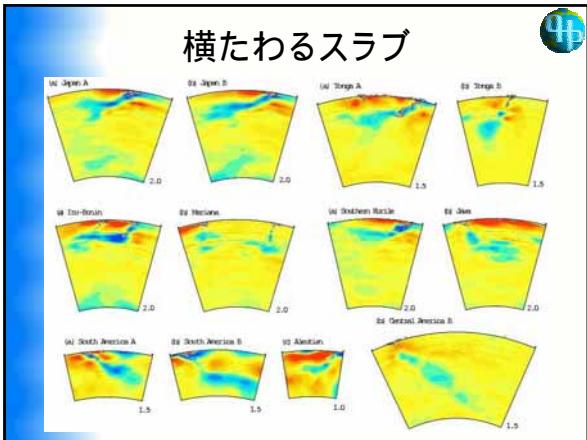
Inoue et al. (1990)



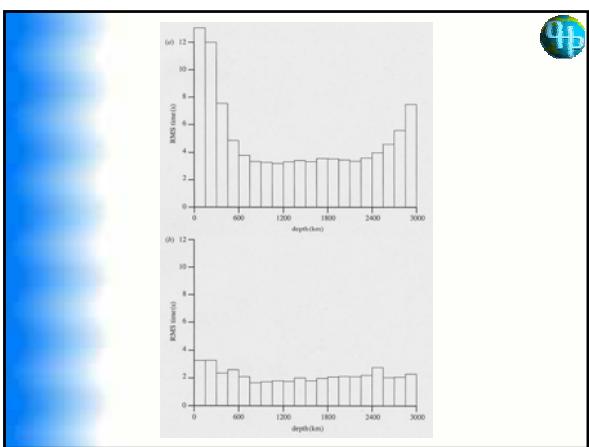
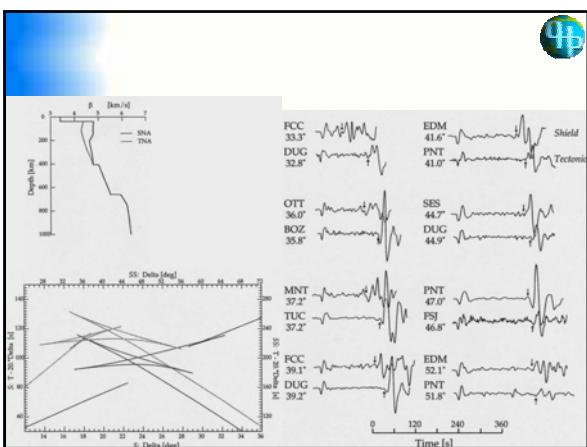
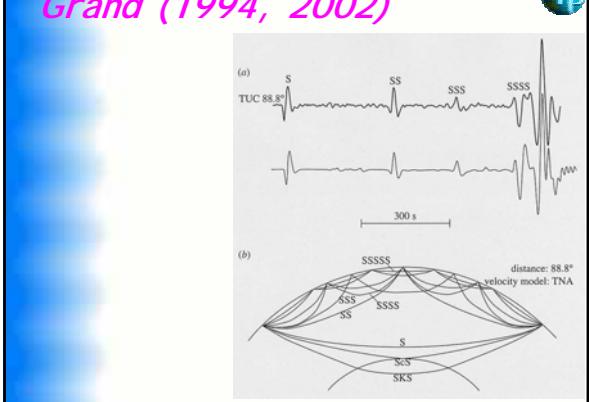
Regularization !

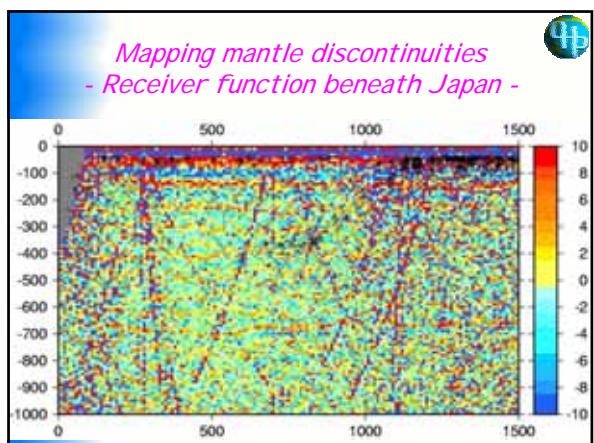
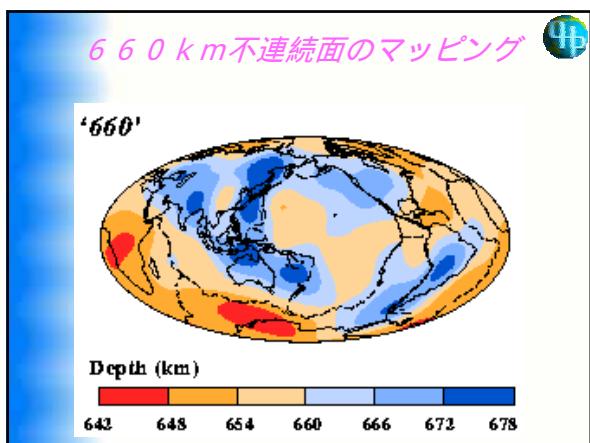
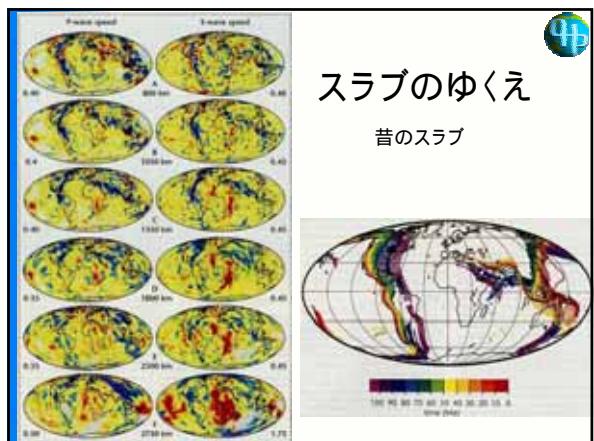
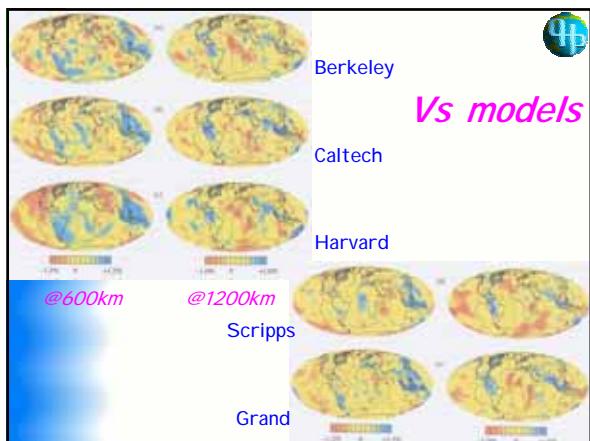
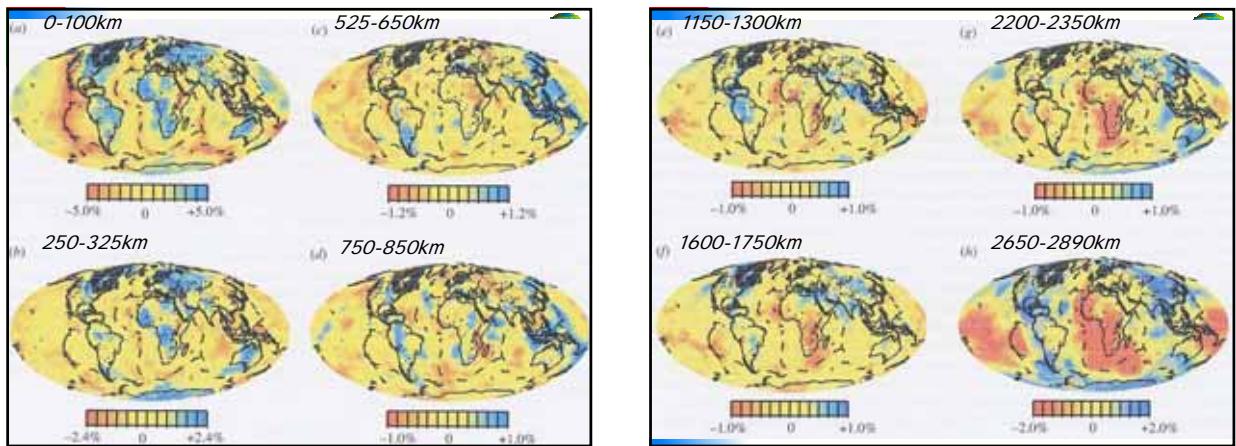


横たわるスラブ



Grand (1994, 2002)





Mantle Tomography -beyond color maps-

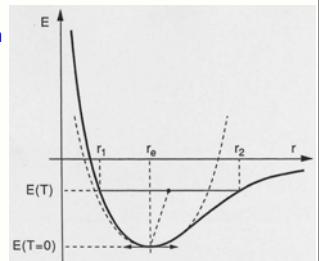
- Karato (1993) *GRL*, 20, 1623-1626
- Masters et al. (2000)
- in "Earth's Deep Interior" (AGU monograph)
- Forte & Mitrovica (2001) *Nature*, 410, 1049-1056
- Davaille (1999) *Nature*, 402, 756-760
- Karato & Kariki (2001) *JGR*, 106, 21771-21783
- Saltzer, Stutzmann & van der Hilst (2004) *JGR*, 109, B06301

scaling: $\delta \ln V_s \sim \delta \ln V_p \sim \delta \ln \rho$?

Thermal effects anharmonicity vs. anelasticity

Karato (1993) *GRL*

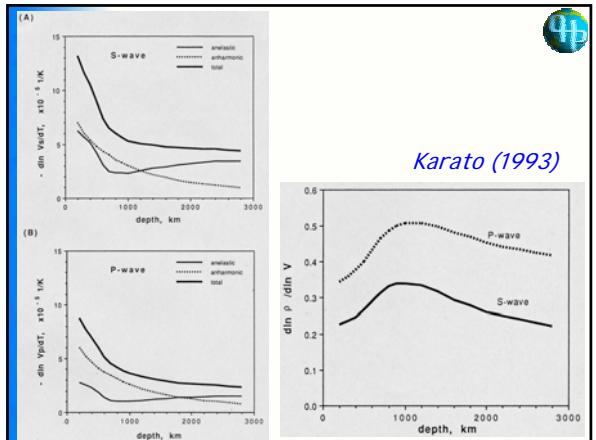
anharmonicity
= thermal expansion



Karato (1993)

For constant Q

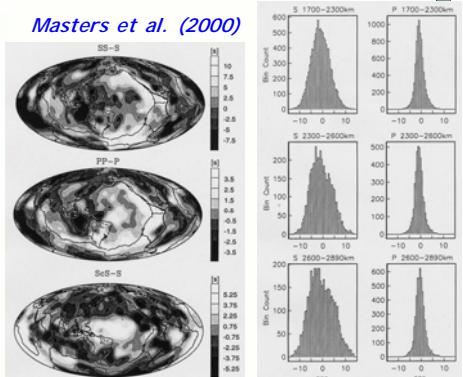
$$\begin{aligned} V(\omega, T) &= V_0(T)(1 + \frac{1}{Q\pi} \ln \omega \tau(T)) \\ \tau(T) &= \tau_0 \exp(H^*/RT) \\ \partial \ln V / \partial T &= \partial \ln V_0 / \partial T - \frac{1}{Q\pi} (H^*/RT^2) \\ (Q^{-1} \ll 1) \end{aligned}$$



Chemical effects?

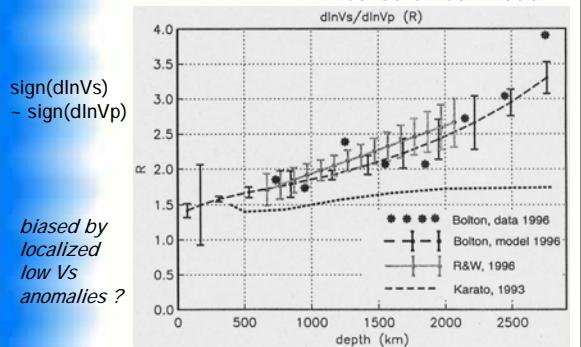
Masters et al. (2000)

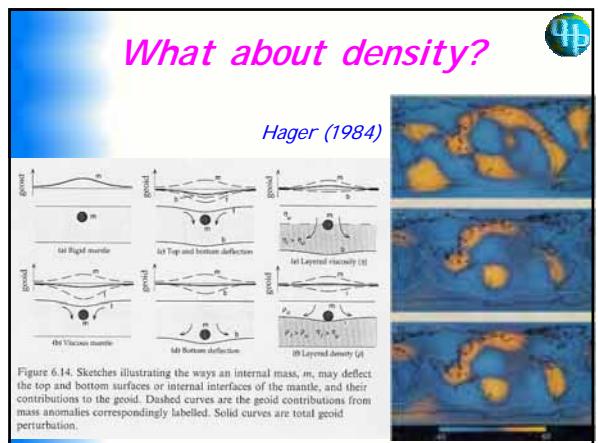
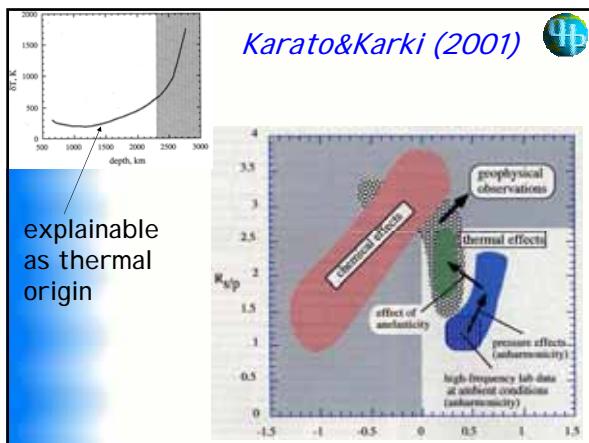
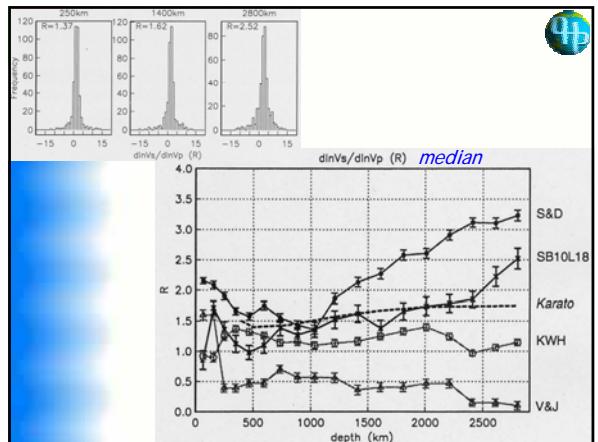
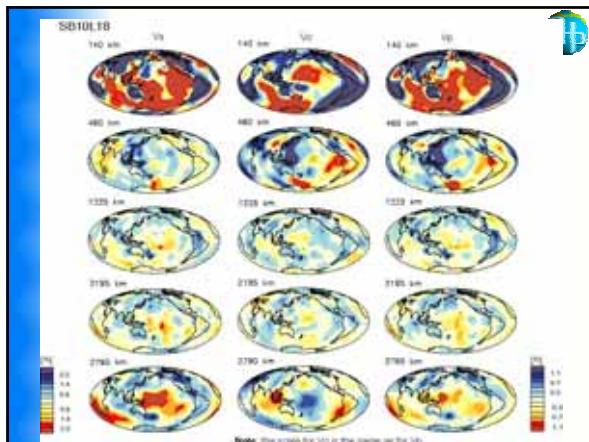
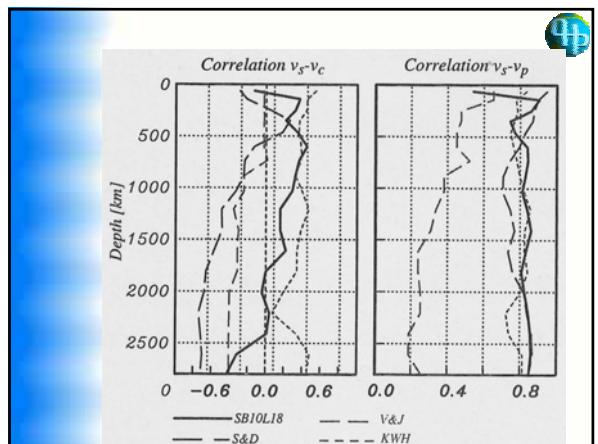
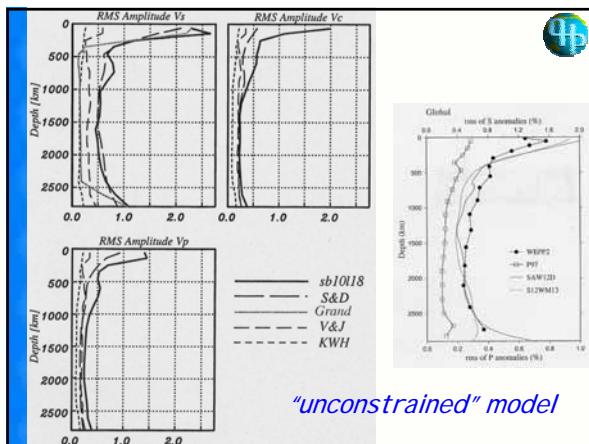
V_p
 vs
 V_s
 vs
 V_c

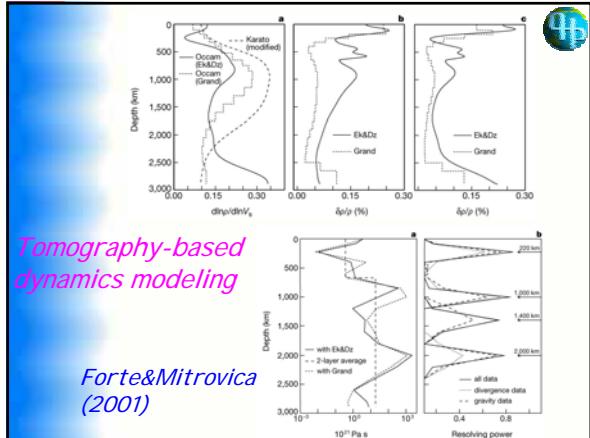


Joint inversion (P & S)

"constrained" model







Forte and Mitrovica (2001)

$$\begin{aligned}\delta \ln V_s &= \frac{\partial \ln V_s}{\partial T} \delta T + \frac{\partial \ln V_s}{\partial X_{Pv}} \delta X_{Pv} + \frac{\partial \ln V_s}{\partial X_{Fe}} \delta X_{Fe} \\ \delta \ln V_\phi &= \frac{\partial \ln V_\phi}{\partial T} \delta T + \frac{\partial \ln V_\phi}{\partial X_{Pv}} \delta X_{Pv} + \frac{\partial \ln V_\phi}{\partial X_{Fe}} \delta X_{Fe} \\ \delta \ln \rho &= \frac{\partial \ln \rho}{\partial T} \delta T + \frac{\partial \ln \rho}{\partial X_{Pv}} \delta X_{Pv} + \frac{\partial \ln \rho}{\partial X_{Fe}} \delta X_{Fe} \\ \delta \ln V_s &= c_{11} \delta T_{eff} + c_{12} \delta X_{eff} \\ \delta \ln V_\phi &= c_{21} \delta T_{eff} + c_{22} \delta X_{eff} \quad (c_{11} \gg c_{21}) \\ \delta X_{eff} &= \delta X_{Pv} + B \delta X_{Fe} \\ \delta T_{eff} &= \delta T + A \delta X_{Fe} \sim \delta T + \epsilon \quad (\epsilon < 10K)\end{aligned}$$

