# The Virtual Muography Institute: Status and Outlook

Cristiano Bozza – University of Salerno Tokyo, Embassy of France 2/10/2017

# Motivations for a virtual institute

- Work together to speed up progress on muography
- Split problems, optimize solutions
- Boost discussion, help prepare common projects
- Break software barriers (different environments/formats)





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# What can VMI do for you -2/2

- Geant4 installed
- Gitlab (software management + build/test/run pipelines)
- Jenkins (continuous integration)

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First project to simulate high energy secondary cosmic rays traversing rock

# Gitlab at VMI – 1

• View source files in web

• Edit «in place» or use local «git» to manage software evolution

• Define build and test procedures

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# Gitlab at VMI –

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### Jenkins at VMI

- Integrated with git
- Custom shell scripts can be called

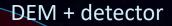
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# MuoGraphSim1 at VMI

- First Geant4 program at VMI
- Simple but flexible
- Decouple cosmic ray simulation from passage through rock and instrument acceptance
- Output files document incoming muon as well as all particles at the detector surface with a proper range (1 m)
- Easy to apply weights to rescale to cosmic ray models

DEM: rectangular mesh (free step and size) Detector: 10m radius sphere, efficiency = 1



Muon generation surface Isotropic flux Spectrum: flat or log-flat (1/E)

# MuoGraphSim1 at VMI

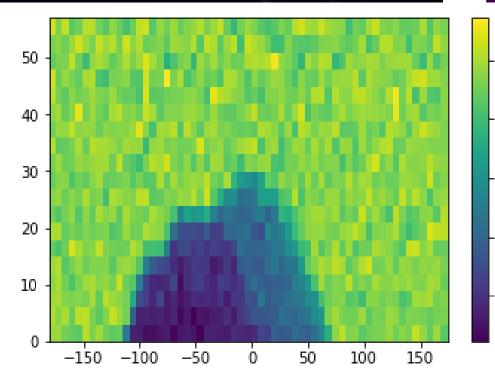
• Example of application to Stromboli

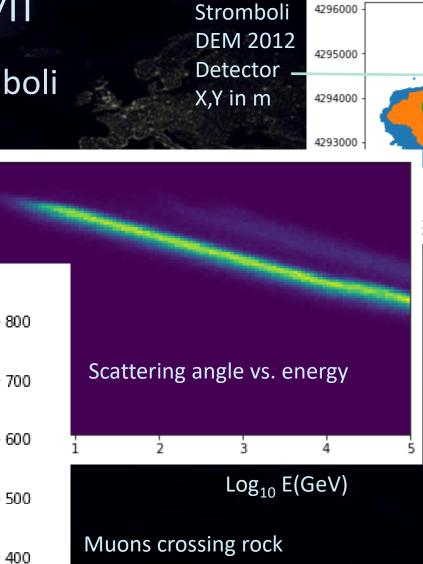
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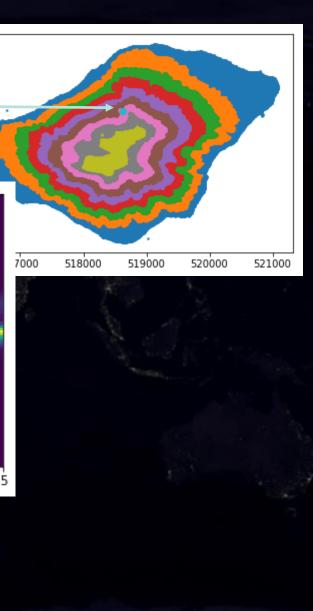
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- 1,5 Mtracks
- 5 hour running





4296000



(Log-flat isotropic spectrum)

### Cloud Hardware at VMI

- CPU: 2×14 cores = 28 real cores, 56 hyperthreaded
- Memory: 64 GiB, possible upgrade to 192 GiB
- Storage: 6 TB
- Upgrades available depending on usage!

# Simulation of secondary cosmic rays – 1/2

- Install CORSIKA or COSMOS at VMI
- Use ready-made CORELib production by KM3NeT for ASTERICS (<u>www.asterics2020.eu</u>), currently distributed through a server in Salerno
- 1% «no Cherenkov» pilot production  $\rightarrow$  550 GB



				Energy range (GeV)	Number of events
High energy model	Low energy model	Option		200-1000	10 <sup>7</sup>
				103-104	10 <sup>7</sup>
		TAULEP	CHARM	104-105	10 <sup>6</sup>
QGSJET01	GHEISHA		Х	105-106	10 <sup>5</sup>
QGSJET01	GHEISHA	Х		106-107	104
QGSJETII- 04	GHEISHA	Х		107-108	10 <sup>3</sup>
EPOS LHC	GHEISHA	Х		108-109	10 <sup>2</sup>

# Simulation of secondary cosmic rays – 2/2

#### Status of CORELib

- Proton-induced showers:
  - HE models : QGSJET01 with CHARM, QGSJET01 with TAULEP, QGSJET-II with TAULEP, EPOSLHC with TAULEP
  - LE model : GHEISHA
  - about 21M Evts per HE model
  - 7 energy bins (2×10<sup>2</sup>GeV-10<sup>3</sup>GeV+equally logarithmically spaced from 1TeV to 10<sup>9</sup>GeV)
  - power-law spectrum with -2 spectral index
  - zenith angle from 0 to 89 degrees
- Nuclei-induced showers:
  - HE model : QGSJET01 with CHARM, QGSJET01 with TAULEP, QGSJET-II with TAULEP, EPOS-LHC with TAULEP
  - LE model : GHEISHA
  - about 21M Evts per HE model
  - 7 energy bins (2×10<sup>2</sup>GeV-10<sup>3</sup>GeV+equally logarithmically spaced from 1TeV to 10<sup>9</sup>GeV)
  - Power-law spectrum with -2 spectral index
  - zenith angle from 0 to 89 degrees

## Conclusions

- VMI is moving its first steps
- Very interesting opportunity for real, tight cooperation
- Geant4 cloud already set up and running
- Possible interaction with other projects
- Capabilities will grow with user demand, so don't hesitate to ask!